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VISSR ATMOSPHERIC SOUNDER

Monthly Progress Report No. 2  
For the period 1 Oct. 1973 to 31 Oct. 1973

Contract No. NAS5-21965

For National Aeronautics and Space Administration  
Goddard Space Flight Center  
Glen Dale Road  
Greenbelt, Maryland 20771

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TABLE OF CONTENTS

I. Introduction . . . . . 3

II. Evaluation of Alternative Detector Configurations. . . . . 3

III. Operating Modes - Dwell Sounding . . . . . 10

IV. Registration Errors - Responsivity Variations. . . . . 13

Addendum: Calculation of Autocovariance Functions

Appendices A, B, C, D, E

## I. Introduction

During the past month and a half discussions with SBRC, GSFC, and NOAA have brought to light new opportunities and new problems which impact the basic VAS detector configuration and VAS operation. Important factors bearing on this situation are: (1) a general consensus exists that  $3.7\mu$  and  $4.3\mu$  channels provide significant improvements in sounding capability; (2) the 6 HgCdTe + 6 InSb detector array suggested by U.W. in September results in a complex dewar package which, according to SBRC, represents substantial risk to the program schedule; (3) previously assumed operational constraints which required imaging/sounding (non-interfering) as a primary mode appeared to be relaxed to the degree that dwell sounding operation could be considered as primary; and (4) greater emphasis was placed by NASA on minimizing multiplexer modifications as much as possible.

As a result of these and other considerations considerable effort during the past month was directed at evaluating possible alternative detector configurations and operating modes. Additional areas of investigation included registration problems and improvements in the large time interval model for the autocovariance function. The latter work is reported in the addendum.

## II. Evaluation of Alternative Detector Configurations

Suggestions of alternate detector configurations were made with a number of objectives in mind, many of which are conflicting. Among the most important are the following:

- (1) minimize the number of detectors to reduce dewar complexity and simplify MUX modifications (reduce total sampling rate required);
- (2) increase the number of detectors to improve capability to sound without

interfering with imaging;

- (3) increase detector size to obtain better radiometric accuracy and minimize detector registration problems;
- (4) reduce detector size to improve cloud discrimination and imaging capabilities and reduce detector registration problems;
- (5) increase total detector area to decrease sounding time; and
- (6) decrease total detector area to minimize dissipation on the cold finger.

In all, seven different detector configurations have been considered in terms of these objectives. These are identified by numbers of detectors of different size and/or type as outlined in Figure 1, which also displays the geometrical arrangements for each possibility. Only the "6" configuration is absent from Figure 1. It is the original configuration specified in the April 1972 document "SMS Sounder Specification" and is the same as the "6 + 6" configuration of Figure 1 except that the InSb array of 6 detectors is omitted. Note that Figure 1 does not accurately portray the detailed east-west positioning of the detector elements which depends on lead runout requirements, sampling intervals, etc. One should also note that the numbers in the configuration I.D. also state the numbers of detectors which must be sampled in any given scan. For example, in the "4 + 2" configuration 4 HgCdTe detectors are sampled during scans which have a filter from any of bands 1-10, while 2 InSb detectors are sampled during other scans.

The configurations were analyzed in terms of a number of performance characteristics and compared in Table 1. Before evaluating these comparisons it is first necessary to explain each of the performance factors listed. These are discussed in the following.

DETECTOR CONFIGURATIONS

IDENTIFICATION

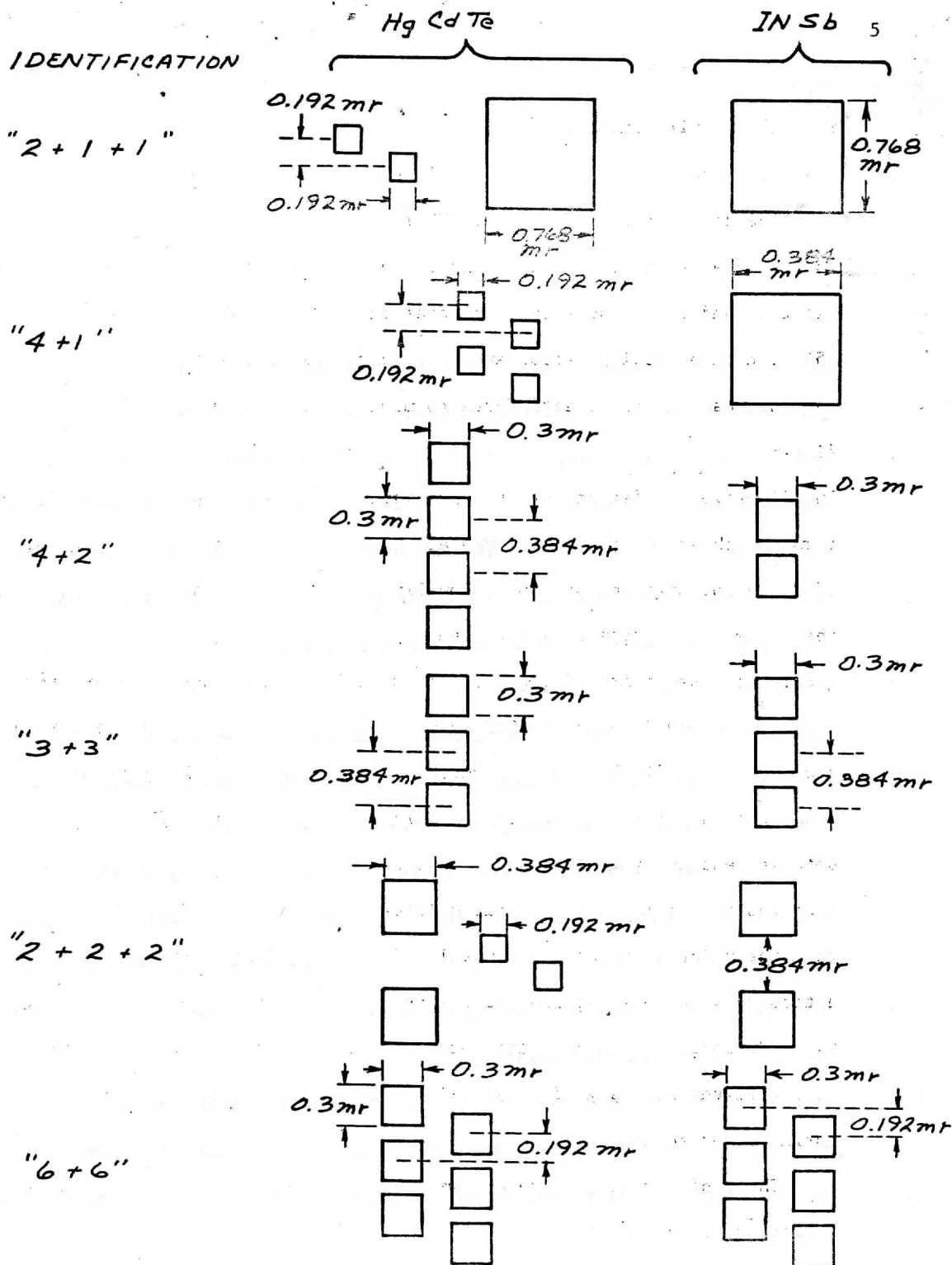


Figure 1 Alternative Detector Configurations and Identification  
 The milliradian dimensions are equivalent to subsatellite FOV's of 6.9 km for 0.192mr, 10.7 km for 0.3 mr, 13.7 km for 0.384mr, and 27.5 km for 0.768mr.

(a) RELATIVE SOUNDING TIME - This is an approximate estimate of the relative time to obtain sufficient radiometric accuracy for sounding for clear conditions. It is based on noise levels for a typical channel, not the weighted characteristics of all channels. Since the effects of clouds are not included, the advantages of high spatial resolution do not appear in this column. Note that the "6 + 6" configuration is defined to have a time factor of unity. Larger numbers indicate proportionately longer sounding times.

(b) DISSIPATION - This is a relative measure of power dissipated on the cold finger by HgCdTe detector joule heating. It does not include lead conduction, aperture effects, and other factors which are not as significant in this comparison. The "6 + 6" is taken to have unit dissipation. The absolute dissipation depends on detector characteristics which vary more than the relative dissipations between configurations. At this time it appears unlikely that any configuration would have to be rejected on the basis of excessive dissipation.

(c) SOUNDING RESOLUTION - This is the angular FOV of the detectors used to measure radiances in the spectral intervals used for sounding (15 $\mu$ , 4.3 $\mu$ , and 6.7 $\mu$  bands). Retrieved soundings would be likely to have lower resolution than this.

(d) IMAGING RESOLUTION - This is the angular FOV of the detectors used to measure radiances in the 11.1 $\mu$  window channel.

(e) SOUNDING SAMPLES/DWELL - This is the number of samples per dwell that could be obtained without changing the multiplexer bit rate. The dwell time used is one typical to the 15 $\mu$  channels (or bands). The "W MOD" column indicates what could be achieved by changing the MUX format (and also omitting sync words); the "W/O MOD" column indicates what can be obtained within

the present MUX format by switching from one detector to the next at the same rate as the present switching between primary and redundant detectors.

(f) NEED FOR LIGHT PIPE - Large detectors, which are more frequently observing cloud contamination, must produce accurate registration of spatial responses from channel to channel. However, non-uniformity of detector responsivity across the detector surface ( $\pm 10\%$  for InSb and greater than  $\pm 50\%$  for HgCdTe) makes this impossible unless a light pipe is used to diffuse radiation from each image point over the entire detector surface. For small detectors, which have a greater tendency to observe either cloud or clear conditions, the responsivity variations do not represent such a great problem. However, no absolute requirement can be used to determine whether or not light pipes must be used since this varies with data processing procedures and depends on how accurately light pipes can be constructed. Consequently the estimates of need are just estimates at this time.

(g) POSSIBILITY OF LIGHT PIPE - This is an estimate by SBRC of the feasibility of constructing a light pipe of the required size with present fabrication techniques. It does not represent an accurate determination of whether the achievable precision meets the ultimate geometrical requirements (as yet uncertain).

(h) RISK FACTOR - These are very approximate estimates of the risk to the program schedule for each detector configuration. These are based on information provided by SBRC. A risk factor of 2 represents an intolerable risk (program delays are quite possible). A risk factor of 1 represents a marginal risk, while a risk factor of 0.25 represents very little risk.

(i) S/C MODIFICATION DIFFICULTY - In order to achieve sampling rates which are acceptable three levels of modification were considered. These are: 1. no MUX or A/D converter modifications, only low speed switching

between detector arrays at line rates within the VISSR package; 2. high speed switching among more than two detectors at the same rate as presently done between two detectors, with switching done on the S/C and no basic changes in MUX format or A/D converters; 3. changes in MUX format, omission of some sync words, A/D modifications as required to increase total sampling rate.

Note that none of these modifications change the transmission bit rate. Nor do they affect the visible sampling and formatting.

(j) IMAGING/SOUNDING CAPABILITIES - This is the number of sounding channels that can achieve full spatial coverage in one pass in the imaging/sounding mode of operation. It is assumed that the 11.1 $\mu$  window channel always requires full spatial coverage in this mode.

Any evaluation of these comparisons must face the difficulty of assigning appropriate weights to each of the factors considered. In many cases this is just a matter of opinion or general consensus rather than an objective logical process. In the present context of this program, however, significant conclusions can be drawn from these comparisons: (1) the "6 + 6" configuration involves an unacceptable risk in terms of dewar package fabrication; (2) the "4 + 1" configuration requires too much sounding time (4 x the minimum); (3) the "6" configuration does not have the proven cloud discriminating capabilities of a 3.7 $\mu$  window channel; and (4) the "2 + 1 + 1" configuration has relatively high dissipation and risk as well as low sounding resolution. Of the remaining configurations which do not have serious disadvantages, or a combination of relatively serious disadvantages, the basic choice is between the "4 + 2" and the "2 + 2 + 2", the former providing somewhat better imaging/sounding capabilities, while the latter provides the lower dewar risk and the lowest impact on S/C modifications of all the configurations.



Table 1. Comparison of Alternative Detector Configurations

DETECTOR ARRAY	RELATIVE SOUNDING TIME	RELATIVE DISSIPATION	SOUNDING RESOLUTION	IMAGING RESOLUTION	SOUNDING SAMPLES/DWELL		LIGHT PIPE	NEED POSSIBLE	RISK FACTOR	S/C MOD DIFFICULTY	IMAGING/SOUNDING CAPABILITIES
					W MOD	W/O MOD					
"2 + 1 + 1"	0.9	1.23	0.768mr	0.192mr	--	18.4	yes	yes?	1.0	1	2
"4 + 1"	3.7	.27	0.192mr <sup>(1)</sup>	0.192mr	1.83	1.15	yes	yes?	1.0	3	3
"4 + 2"	1.5	.67	0.30mr	0.30mr	2.9	1.8	?	?	0.7	2-3	3
"3 + 3"	2.0	.50	0.30mr	0.30mr	3.6	2.4	?	?	1.0	2	2
"2 + 2 + 2"	1.8	.68	0.384mr	0.192mr	---	4.6	?	?	0.7	1	2
"6 + 6"	1	1.0	0.30mr	0.30mr	2.0	1.2	?	?	2.0	3	5
"6"	1 + Q <sup>(2)</sup>	1.0	0.30mr	0.30mr	2.0	1.2	?	?	0.3	3	5

NOTES: (1) The 0.192mr applies for 15 channels, 0.768mr at 4 .

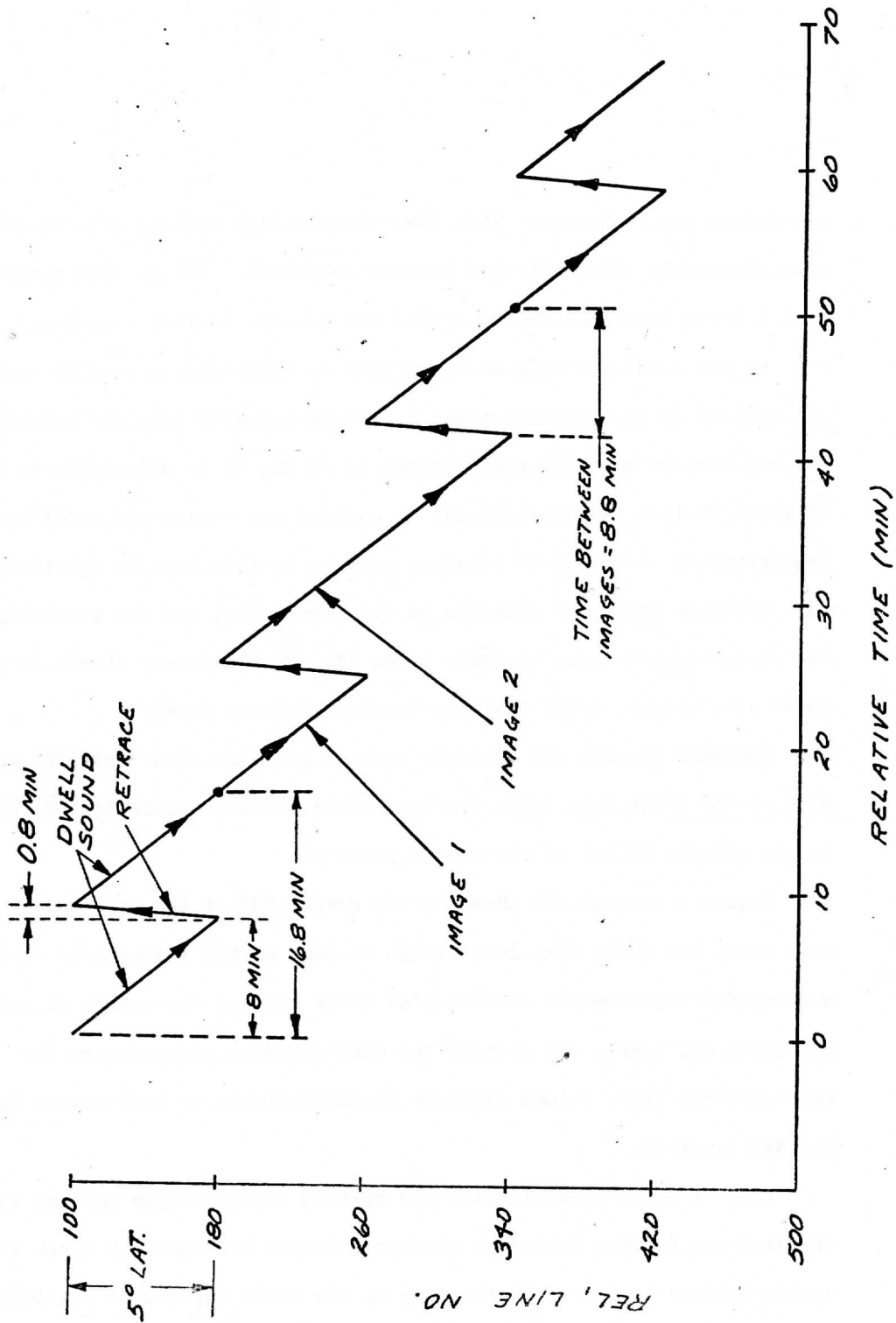
(2) Time allotted for Q bands would substantially increase the total time unless external data used.

On the whole, the "2 + 2 + 2" configuration provides dwell sounding capabilities (considering clear sounding time and cloud discrimination effectiveness) as good as any of the configurations, and also has the best chance of success (highest feasibility) of any of those configurations including InSb detectors.

### III. Operating Modes - Dwell Sounding

During dwell sounding (with the "2 + 2 + 2" configuration, for example) it is possible to meet NEN requirements at a spatial resolution of 30 km while scanning one IR line per 20 spins (average). This amounts to about 20° of latitude (near the subsatellite point) per hour. Since the value of these soundings would be greatly augmented by simultaneous wind field information some modification of the dwell sounding mode should be considered. Instead of merely scanning down 30° of latitude to obtain sounding coverage it would be advantageous to break the 30° swath into smaller pieces (say 5° to 10° swaths) and scan each twice at twice the rate to obtain an image pair which could be used to derive cloud displacements ("winds"). Actually this process would produce 13 image pairs (12IR and 1 visible) which would provide a unique opportunity to determine the vertical structure of the wind field.

In simplified terms the proposed dwell sounding mode would operate as follows: (1) scan down from start line to end line at an average rate of 1 line every 10 spins; (2) retrace at 1 line per spin from end line to start line; (3) scan down again at 1 line per ten spins. At this point a new subframe could be initiated. An example of a dwell sounding sequence of 5° subframes is presented in Figure 2. The image time separation of 8.8 min. may not be optimum considering the geometrical precision of the VISSR scanner and the required wind accuracy. A 10° subframe with 17.4 min. image



separation may be better. This time optimization problem will be addressed more thoroughly after SMS data becomes available. (It is also possible that a three image set may be needed for quality control purposes.)

If the dwell sounding mode operates as suggested, a problem arises as to the details of the scanning process. If 20 spins/IR line (80 spins/dwell) are required to meet NEN requirements at 30 km, is it acceptable to obtain 10 spins/IR line (40 spins/dwell) on each of two passes and still meet requirements? A number of possible ways to do this must be considered:

- (a) Sequence through 6 channels on the first pass, and the remaining 6 on the second pass (window channels 8 and 12, which use very little of the total spin budget, would actually be used on every pass);
- (b) Sequence through all channels on each pass but cover only 50% of the area on the first pass (with missing lines) and the remaining 50% (filling in the missing lines) on the second pass; or
- (c) Sequence through all channels and cover 100% of the scanned area on each pass, but dwell only long enough to come within a factor of  $\sqrt{2}$  of meeting NEN requirements on each pass (thus halving the number of spins per dwell per pass), and average the corresponding measurements (or corresponding clear column radiance determinations) on both passes to obtain the NEN required.

Each of these possibilities has certain disadvantages related to registration (slight motion of weather features and apparent earth motion occurs between passes). Whether any of the three approaches can duplicate the performance obtained in a single slow scan (20 spins/IR line average) is currently a matter of conjecture needing further study. If closer examination dictates, additional methods for obtaining image pairs will be considered.

#### IV. Registration Errors - Responsivity Variations

In processing cloud contaminated sounding channel data to obtain equivalent clear column radiances it is usually necessary to use paired (geometrically adjacent) fields of view and the 11.2 $\mu$  window channel in the following formula

$$I_{c,v} = \frac{I_v^{(1)} - N^* I_v^{(2)}}{1 - N^*} \quad (1)$$

where

$$N^* = \frac{I_w^{(1)} - I_{c,w}}{I_w^{(2)} - I_{c,w}} \quad (2)$$

and we use the definitions

$I_{c,v}$  = clear column spectral radiance for spectral band  $v$

$I_{c,w}$  = clear column spectral radiance for a window channel

$I_v^{(j)}$  = observed radiance from FOV #j in band  $v$

$I_w^{(j)}$  = observed radiance from FOV #j in a window channel

In order to apply this formula it is necessary to know  $I_{c,w}$  and to have the following conditions satisfied:

- (1)  $N^* \neq 1.0$
- (2) clouds in FOV's 1 and 2 have grey emittance-transmittance properties
- (3) the height of the clouds is the same in FOV's 1 and 2

It is the function of the 3.7 $\mu$  window channel to provide a test for conditions (2) and (3). For example, if equation (2) is applied for both 3.7 $\mu$  and 11 $\mu$  windows and the two  $N^*$  values determined do not agree then the assumed cloud conditions are not present. In order to make this check, however, it is essential that both 3.7 $\mu$  and 11.2 $\mu$  windows have identical FOV's, i.e. that they observe the same thing. Otherwise the check is meaningless.

Registration of 3.7 $\mu$  and 11.2 $\mu$  window channels is complicated by the fact that neither detector has uniform responsivity over the detector surface.

The HgCdTe detectors typically have (according to SBRC) 50% responsivity variations, while the InSb detectors (for the  $3.7\mu$  window) have 10% responsivity variations. Thus, if two detectors of the same size are considered, these variations of sensitivity in the projected FOV are so large as to make the paired FOV corrections hopeless. However, if small HgCdTe detectors (0.192mr FOV) are used in conjunction with large InSb detectors (0.384mr FOV); it is possible that, over the larger FOV, the assemblage of small HgCdTe FOV's may have an effective responsivity variation significantly smaller than that of each individual small FOV.

An indication of this improvement can be seen from simulation results from somewhat different test conditions. Four lines and six samples per line of 0.3mr FOV's were averaged to obtain radiances over a 30 km square. Each 0.3mr FOV was broken into 9 subsegments with individual weights simulating 50% responsivity variations. Sloped and Peaked (S and P) weight distributions were used with maxima at Bottom, Right, or Center (B, R or S) of the detector (hence the notation BS for bottom sloped). A variety of weights and a variety of seven cloud distributions were used in the test. Results were compared to the same average with flat field detectors. As indicated in Figure 3 resulting radiometric errors depend on cloud distribution, weight distribution, and the difference between cloud and clear radiances. For approximately 50% cloud cover the RMS overall error is approximately 2% of the difference. For many cases this means errors of the order of  $1 \text{ erg}/(\text{sec-cm}^2\text{-ster-cm}^{-1})$ , a quite unsatisfactory result. Unless better responsivity uniformity can be obtained larger averaging areas, smaller cloud amounts, or warmer clouds, or all three restrictions must be applied to obtain adequate radiometric registration.

SCENE:	(1)	(2)	(3)	(4)	(5)	(6)	(7)	RMS %	
1) BS, RS, BP, CP								2.8%	1.3
2) BS, RS, BS, RS	3.7%	1.9%	0.0%	0.9%	-1.5%	1.4%	-0.9%	1.8	
3) BP, CP, BP, CP	1.8%	0.0%	-1.7%	0.9%	0.7%	2.7%	0.7%	1.2	
4) BS, BS, BS, BS	3.2%	0.0%	4.2%	0.5%	0.0%	-0.5%	-0.2%	2.0	
5) RS, RS, RS, RS	3.4%	3.7%	0.0%	-0.5%	0.3%	-1.2%	0.3%	2.0	
	3.1	1.9	2.0	0.7	0.9	1.6	0.6	1.8	

Figure 3 RADIOMETRIC ERROR =  $-\Delta \times$  (tabulated value)

$\Delta$  = radiance difference between clear and cloudy conditions

Example:  $\Delta = 50$  ergs/etc., Detector wts. 1), Scene (6)  
 Radiance error (compared to flat weights) =  $50 \times \frac{1.5}{100} = 0.75$  erg/etc.

It should be noted that there are many other sources of registration errors which have not been discussed, and which are being studied.



CALCULATION OF AUTOCOVARANCE FUNCTIONS

Prepared by Subhash Dandage

[1] Previous work

Until now, the calculation of the autocovariance function,

$$\underline{C(\tau)} = \int_0^{\infty} T^2(f) p(f) \cos 2\pi f \tau df \quad (1)$$

was based on the assumption of an ideal bandpass filter. In other words it was assumed that

$$T^2(f) = \begin{cases} 1 & \text{for } f_{\min} \leq f \leq f_{\max} \\ 0 & \text{for } f < f_{\min} \text{ and } f > f_{\max} \end{cases}$$

also 
$$p(f) = \left(1 + \frac{f}{c}\right) k \quad \text{where } k = \frac{\sigma^2}{\Delta f_N}$$

In such a case  $\frac{C(\tau)}{\sigma^2}$  is given by

$$\frac{C(\tau)}{\sigma^2} = \frac{1}{\Delta f_N} \int_{f_{\min}}^{f_{\max}} \left(1 + \frac{f}{c}\right) \cos(2\pi f \tau) df \quad (2)$$

where  $\Delta f_N = f_{\max} - f_{\min} + f_c \ln \frac{f_{\max}}{f_{\min}}$

Program "COVARM" written in Oct. 71 was used to compute  $\frac{C(\tau)}{\sigma^2}$  for such a case. Use of math tables was done in computation of integrals.

$\frac{C(\tau)}{\sigma^2}$  was also modelled by an equation of the form:

$$\frac{C(\tau)}{\sigma^2} = a - b \ln \tau \quad (3)$$

for the case under consideration, a and b can be computed with the help of the formulae:  $b = \frac{f_c}{\Delta f_N}$ ;  $a = b(-\ln 2\pi f_{\min} - \gamma)$ ,  $\gamma = .577216$  (4)

Refer to the report on program "COVARM" dated 23 Oct. 71 for more details about this special case.

[2] Recent work

The ideal bandpass filter assumption is not realistic. The real filter transfer function should include:

- 1) the presampling filter and
- 2) Restore circuit noise transfer function.

The following parameters are known:

$$f_{\max} = 17450 \text{ Hz}$$

$$f_{\min} = .026 \text{ Hz}$$

$$f_c = 750 \text{ or } 10 \text{ Hz}$$

Presampling filter:

The pole locations for this filter (normalized so that -3 dB radian frequency = 1 rad/sec) are given as

$$p_1 = -0.70557 + j0$$

$$p_{2,3} = -0.67745 \pm j0.94007$$

$$p_{4,5} = -0.54117 \pm j1.82565$$

Program 'FILTER', listed in Appendix A, uses these pole locations to compute the normalized power transfer function for the presampling filter. This power transfer function is plotted in Figure 1.

D.C. Restore noise transfer function:

This transfer function is represented by the following relation

$$N(f) = \sqrt{1 - \frac{2\sin(2\pi f T_1)}{2\pi f T_1 [1 + (2\pi f T_2)^2]^{1/2}} + \frac{1}{[1 + (2\pi f T_2)^2]}} \quad (5)$$

and is plotted in Figure 2.

In this work, it was assumed that  $T_1 = T_2 = 0.3 \text{ sec}$ .

Figure 3 (a,b) show the combined effect of the presampling filter and D.C. Restore. It can be noticed that the oscillations in the D.C.R. noise transfer function cause a small bump in the total transfer function.

[3] Calculations without D.C.R.

The transfer function of the presampling filter was approximated by a sum of 25 transfer functions, each corresponding to an ideal bandpass filter and having a weighting factor of 0.4. This approximation is shown in Figure 1

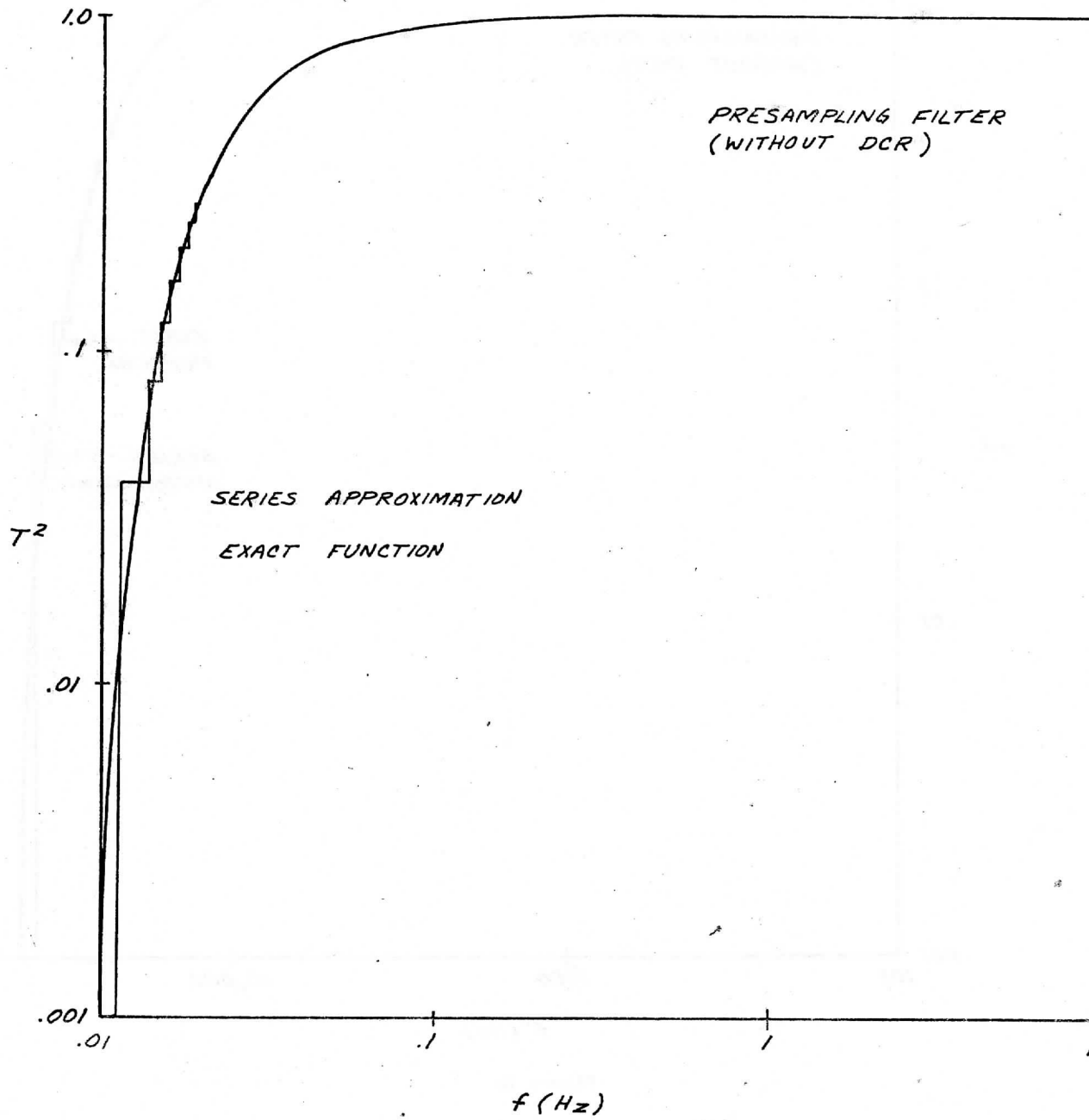


Figure 1a

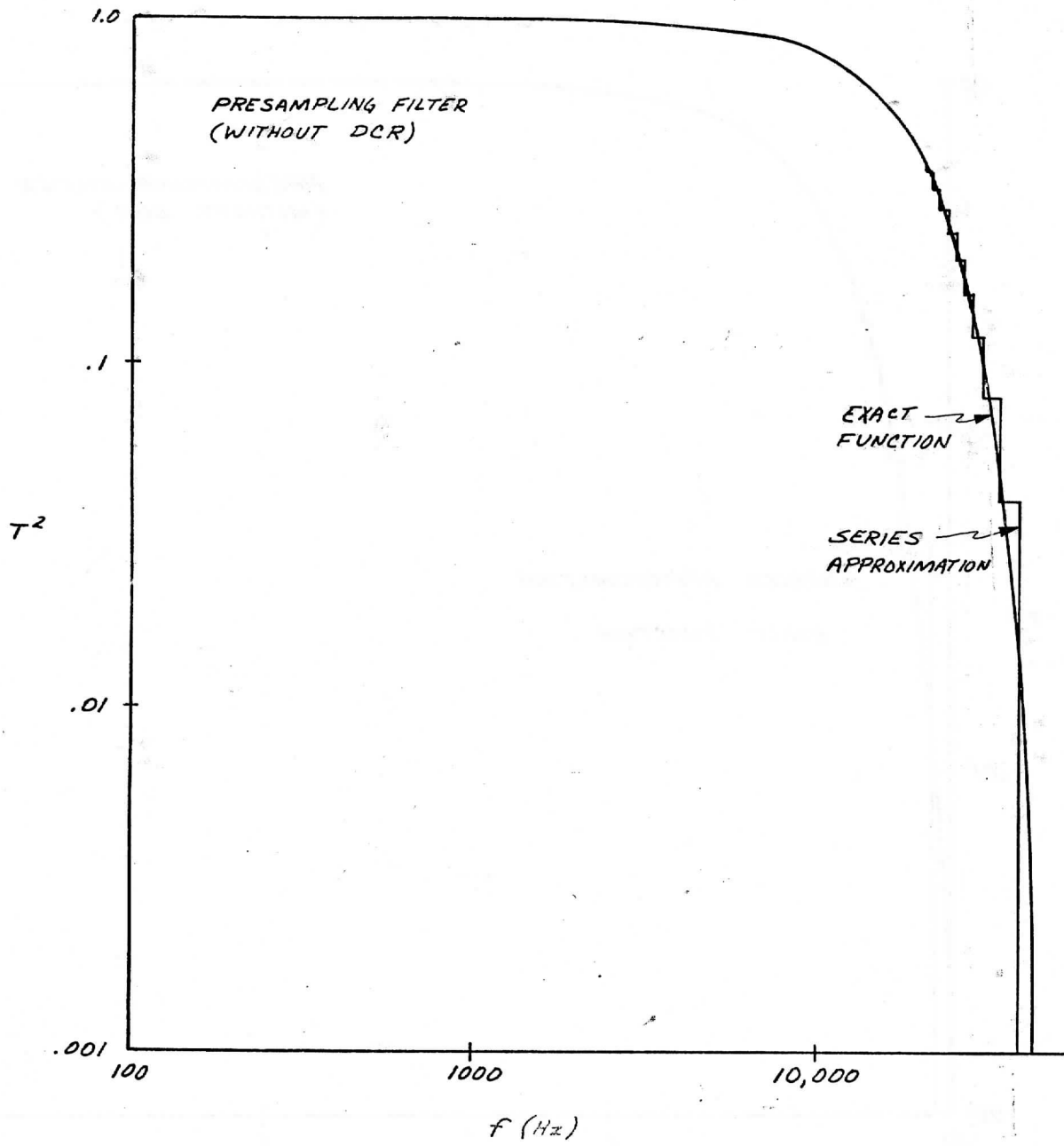


Figure 1b

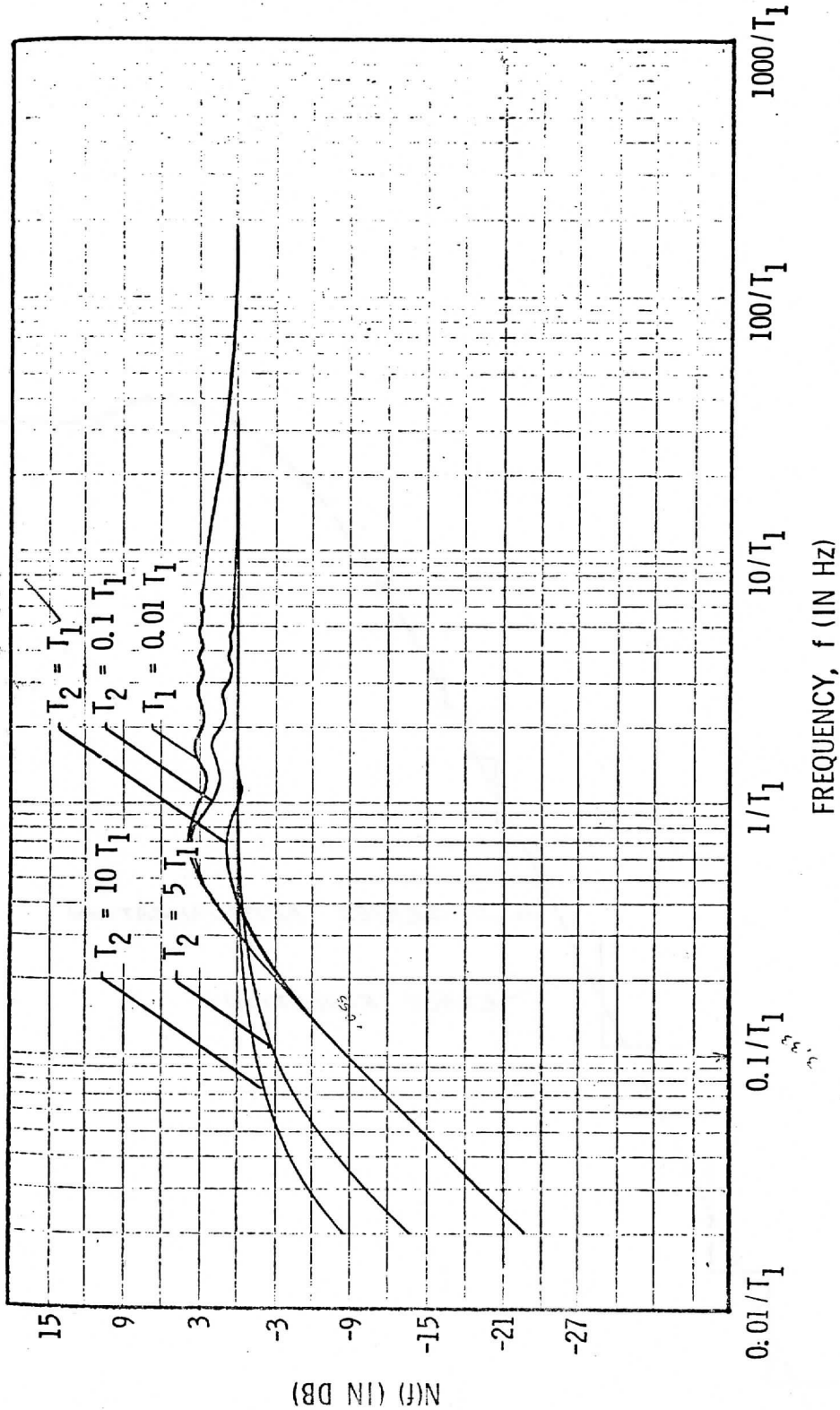


Figure 2

### RESTORE CIRCUIT NOISE TRANSFER FUNCTION

Provided by Santa Barbara Research Corporation,  
 VISSR Sounder Program Initiation meeting at GSFC, 10 July 1973.

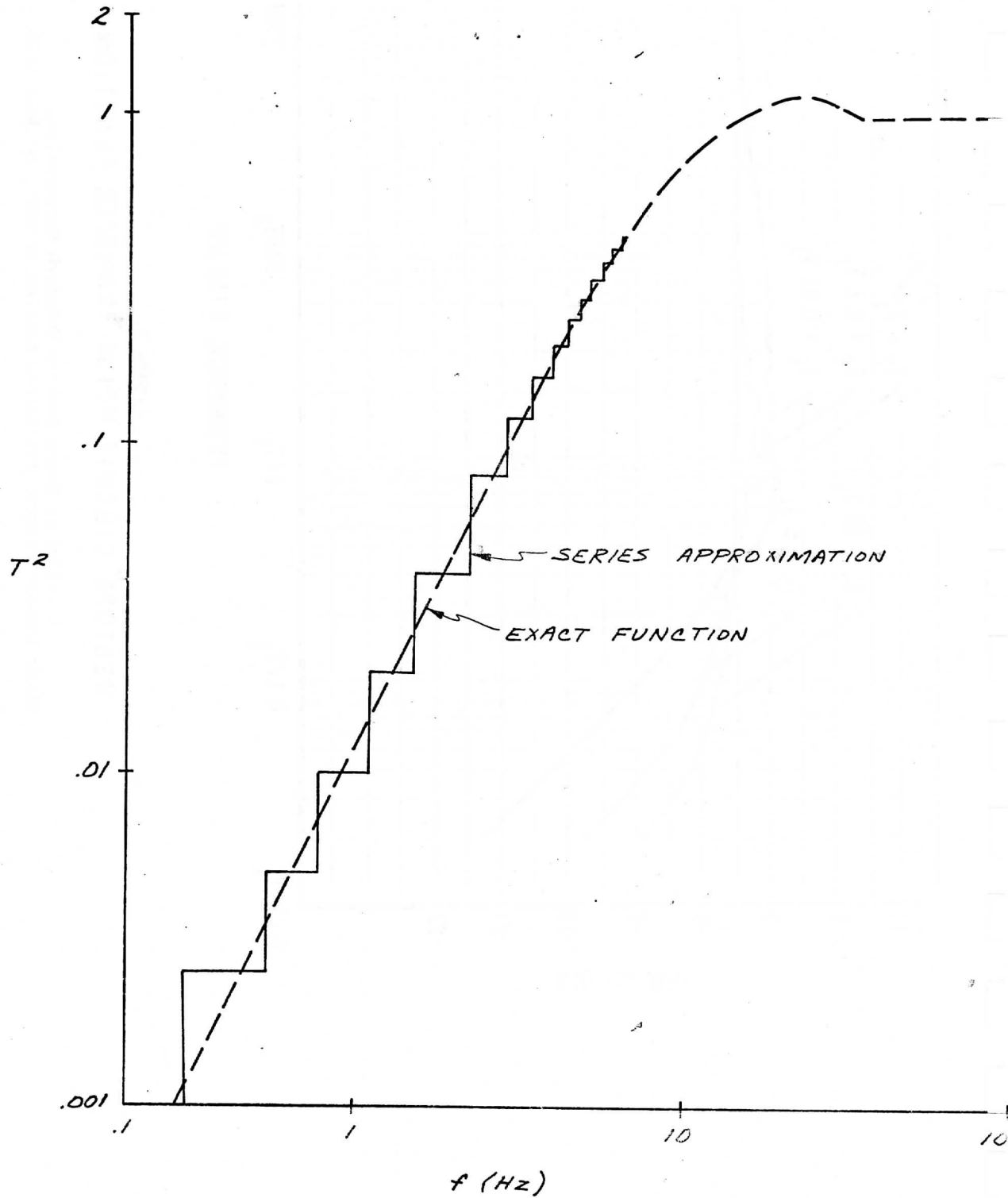


Figure 3a

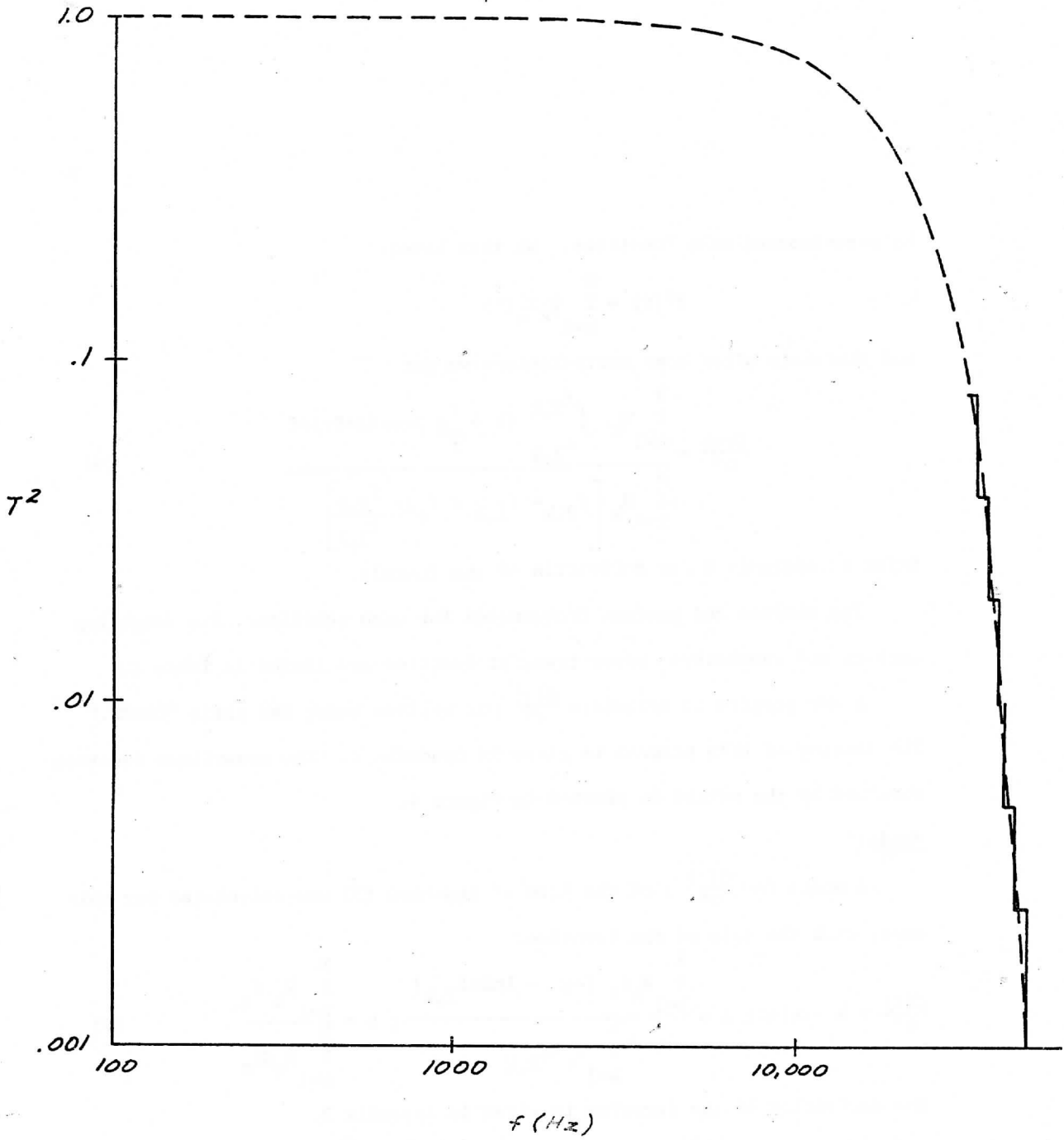


Figure 3b

by superimposed step functions. We then have:

$$T^2(f) = \sum_{k=1}^N W_k T_k^2(f)$$

and therefore after some manipulations we get

$$\frac{C(\tau)}{\sigma^2} = \frac{\sum_{k=1}^N W_k \int_{f_{1,k}}^{f_{2,k}} (1 + \frac{f_c}{f}) \cos(2\pi f \tau) df}{\sum_{k=1}^N W_k \left[ f_{2,k} - f_{1,k} + f_c \ln \frac{f_{2,k}}{f_{1,k}} \right]} \quad (6)$$

Refer to appendix B for derivation of the formula.

The minimum and maximum frequencies for each subfilter, its weighting factors and cumulative power transfer function are listed in Table 1.

A new program to calculate  $\frac{C(\tau)}{\sigma^2}$  was written under the title "CONEW." The listing of this program is given in Appendix C. The covariance function obtained by the method is plotted in Figure 4.

\*Model:

A model for  $\frac{C(\tau)}{\sigma^2}$ , of the form of equation (3) was calculated for this case, with the help of the formulae:

$$\frac{C(\tau)}{\sigma^2} = a - b \ln \tau; \quad a = \frac{\sum_{k=1}^N W_k f_c [-\gamma - \ln 2\pi f_{2,k}]}{N \sum_{k=1}^N W_k \Delta f_{N,k}}; \quad b = \frac{\sum_{k=1}^N W_k f_c}{N \sum_{k=1}^N W_k \Delta f_{N,k}} \quad (7)$$

The derivation of the formulae is given in Appendix B.

The following models were obtained:

$$\begin{aligned} \text{for } f_c = 750 \text{ Hz:} & \quad \frac{C(\tau)}{\sigma^2} = 0.0287 - 0.0272 \ln \tau \\ \text{for } f_c = 10 \text{ Hz:} & \quad \frac{C(\tau)}{\sigma^2} = 0.0006 - 0.0006 \ln \tau \end{aligned} \quad (8)$$

The straight lines superimposing the actual  $\frac{C(\tau)}{\sigma^2}$  curves in Figure 4, represent these models. For large  $\tau$  there seems to be a very good agreement between the models and the accurate calculations.

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\*Program 'CTAU' which computes these models, is listed in Appendix D.



Table 1 Presampling Filter

k	fmin <sub>,k</sub>	fmax <sub>,k</sub>	W <sub>k</sub>	T <sub>k</sub> <sup>2</sup>
1	.35	19,002	.04	1.00
2	.105	4,000	.04	.96
3	.079	5,900	.04	.92
4	.064	7,300	.04	.88
5	.0535	8,500	.04	.84
6	.0475	9,800	.04	.80
7	.043	11,000	.04	.76
8	.038	12,000	.04	.72
9	.034	13,200	.04	.68
10	.0316	14,200	.04	.64
11	.0296	15,400	.04	.60
12	.0278	16,500	.04	.56
13	.026	17,500	.04	.52
14	.0248	18,500	.04	.48
15	.0235	19,300	.04	.44
16	.0222	20,200	.04	.40
17	.021	21,500	.04	.36
18	.0202	22,500	.04	.32
19	.0192	23,750	.04	.28
20	.0182	25,000	.04	.24
21	.0169	26,500	.04	.20
22	.0158	28,000	.04	.16
23	.0149	30,300	.04	.12
24	.0137	33,900	.04	.08
25	.0113	38,500	.04	.04

Fmin = .026  
 Fmax = 17,450  
 F<sub>c</sub> = 750 or 10 Hz

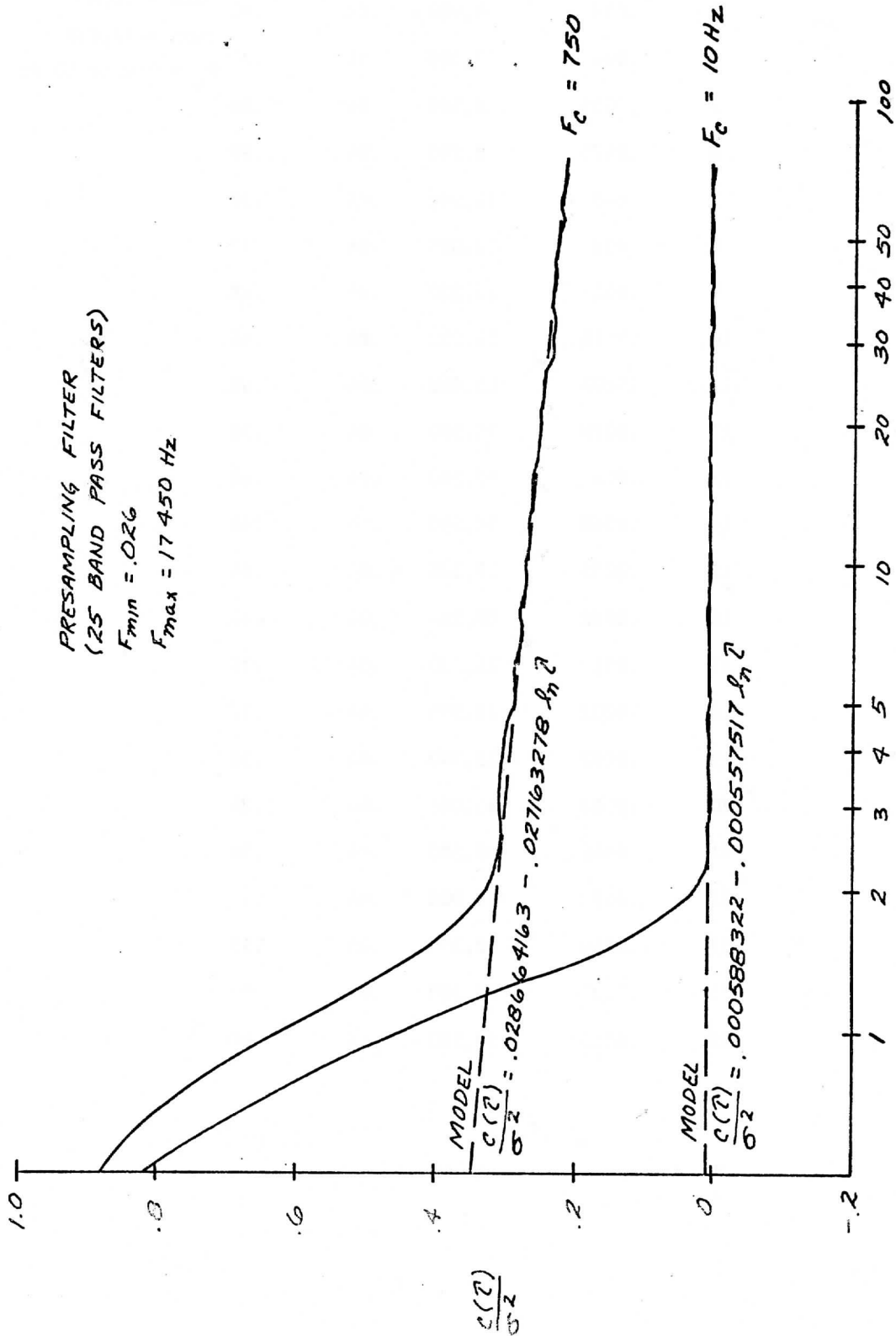


Figure 4

Variance and standard deviation of  $\frac{C(\tau)}{\sigma^2}$

Program 'RTSPAC' written by Ralph Dedecker calculates the variance and standard deviation of  $\frac{C(\tau)}{\sigma^2}$ , for a given  $\frac{C(\tau)}{\sigma^2}$  array. The results obtained are given in Table 2.

To include the effect of D.C.R. in the variance calculations, the following values of 'DCR' were used

$$\text{For } f_c = 750 \text{ Hz} \quad \text{DCR} = 0.0908$$

$$f_c = 10 \text{ Hz} \quad \text{DCR} = 0.00189$$

[4] Calculations with D.C. Restore

The method for analysis of this case is the same as the one in [3]. As shown in Figure 3 (a,b), in this case 32 bandpass filters approximate the effective transfer function. Three of these filters (forming the bump) have very short bandwidths. The weighting factors for different filters are different in this case, as can be seen from Table 3.

Formula 6 (Program CONEW) is used to compute the covariance functions plotted in Figure 5. Also superimposed in the figure, are the following models obtained from program 'CTAU'.

$$\text{For } f_c = 750 \text{ Hz:} \quad \frac{C(\tau)}{\sigma^2} = -.1544 - .03541n\tau$$

$$\text{For } f_c = 10 \text{ Hz:} \quad \frac{C(\tau)}{\sigma^2} = -.0027 - .00061n\tau$$

The variance and standard deviation results for this case are given in Table 4.

The deviation between the model and the calculations in case of  $f_c = 750$  Hz is found to be a result of the bump. As mentioned in Appendix B, the model is good for  $2\pi f_{\max} \tau > 1$ . In case of the 3 filters forming the bump, this is not true. Therefore it was decided to repeat the calculations neglecting the bump.

Table 2

MPLE NO	KM	VARIANCE OF MEAN
1	.53700000+01	.10000000+01
2	.10740000+02	.31641003+01
3	.16110000+02	.58524415+01
4	.21480000+02	.90091065+01
5	.26850000+02	.12631961+02
6	.32220000+02	.16688560+02
7	.37590000+02	.21174129+02
8	.42960000+02	.26072104+02
9	.48330000+02	.31381825+02
10	.53700000+02	.37094728+02
11	.59070000+02	.43198096+02
12	.64440000+02	.49693185+02
13	.69809999+02	.56569643+02
14	.75179999+02	.63829382+02
15	.80549999+02	.71457460+02
16	.85920000+02	.79458282+02
17	.91290000+02	.87820411+02
18	.96660000+02	.96546459+02
19	.10203000+03	.10562528+03
20	.10740000+03	.11506239+03
21	.11277000+03	.12484785+03
22	.11814000+03	.13498357+03
23	.12351000+03	.14546313+03
24	.12888000+03	.15628614+03
25	.13425000+03	.16744803+03
26	.13962000+03	.17894458+03
27	.14499000+03	.19077849+03
28	.15036000+03	.20294269+03
29	.15573000+03	.21543603+03
30	.16110000+03	.22825502+03
31	.16647000+03	.24137782+03
32	.17184000+03	.25486302+03
33	.17721000+03	.26864648+03
34	.18258000+03	.28275138+03
35	.18795000+03	.29716931+03
36	.19332000+03	.31190559+03
37	.19869000+03	.32695605+03
38	.20406000+03	.34232262+03
39	.20943000+03	.35800462+03
40	.21480000+03	.37399792+03
41	.22017000+03	.39030497+03
42	.22554000+03	.40691892+03
43	.23091000+03	.42384353+03
44	.23628000+03	.44107407+03
45	.24165000+03	.45860802+03
46	.24702000+03	.47644397+03
47	.25239000+03	.49458153+03
48	.25776000+03	.51301531+03
49	.26313000+03	.53174882+03
50	.26850000+03	.55077829+03
51	.27387000+03	.57010579+03
52	.27924000+03	.58973232+03
53	.28461000+03	.60965230+03
54	.28998000+03	.62986642+03

WITHOUT D. C. RESTORE

 $F_C = 750 \text{ Hz}$ 

DCR = 0.0908

55	.29535000+03	.65037558+03
56	.30072000+03	.67116869+03
57	.30609000+03	.69225254+03
58	.31146000+03	.71362164+03
59	.31683000+03	.73527635+03
60	.32220000+03	.75722037+03
61	.32757000+03	.77945550+03
62	.33294000+03	.80197333+03
63	.33831000+03	.82477165+03
64	.34368000+03	.84784839+03
65	.34905000+03	.87120355+03
66	.35442000+03	.89484123+03
67	.35979000+03	.91875958+03
68	.36516000+03	.94295110+03

-----  
 LINES NO SAMS AREA(KMXKM) VARIANCE STD DEV  
 -----

1	2	10.74	10.74	.79103	.88940
2	3	17.61	16.11	.32514	.57021
3	5	24.49	26.85	.16843	.41040
4	6	31.36	32.22	.11589	.34043
5	7	38.23	37.59	.08643	.29398
6	8	45.11	42.96	.06790	.26057
7	10	51.98	53.70	.05299	.23020
8	11	58.86	59.07	.04463	.21125
9	12	65.73	64.44	.03834	.19581
10	14	72.60	75.18	.03257	.18046
11	15	79.48	80.55	.02897	.16992
12	16	86.35	85.92	.02587	.16083
13	17	93.22	91.29	.02338	.15289
14	19	100.10	102.03	.02090	.14457
15	20	106.97	107.40	.01918	.13848
16	21	113.84	112.77	.01769	.13302
17	22	120.72	118.14	.01641	.12808
18	24	127.59	128.88	.01507	.12278
19	25	134.46	134.25	.01410	.11875
20	26	141.34	139.62	.01324	.11505
21	28	148.21	150.36	.01233	.11102
22	29	155.09	155.73	.01164	.10791
23	30	161.96	161.10	.01103	.10501
24	31	168.83	166.47	.01047	.10231
25	33	175.71	177.21	.00987	.09934
26	34	182.58	182.58	.00941	.09699
27	35	189.45	187.95	.00898	.09479
28	37	196.33	198.69	.00853	.09236
29	38	203.20	204.06	.00817	.09041
30	39	210.07	209.43	.00785	.08858
31	40	216.95	214.80	.00754	.08683
32	42	223.82	225.54	.00721	.08490
33	43	230.70	230.91	.00695	.08334
34	44	237.57	236.28	.00670	.08186
35	46	244.44	247.02	.00643	.08021
36	47	251.32	252.39	.00622	.07886
37	48	258.19	257.76	.00602	.07756
38	49	265.06	263.13	.00583	.07634
39	51	271.94	273.87	.00562	.07497

40	52278.81279.24	.00545	.07384
41	53285.68284.61	.00529	.07276
42	54292.56289.98	.00514	.07171
43	56279.44300.72	.00498	.07055
44	57306.30306.09	.00484	.06949
45	58313.18311.45	.00471	.06843
46	60320.05322.20	.00457	.06737
47	61326.93327.57	.00445	.06631
48	62333.80332.94	.00435	.06525
49	63340.67338.31	.00424	.06419
50	65347.55349.05	.00412	.06313

~~508 .00389~~

MPLE NO. KM. VARIANCE OF MEAN

1	.53700000+01	.10000000+01	WITHOUT D. C. RESTORE
2	.10740000+02	.28808247+01	$F_C = 750 \text{ Hz}$
3	.16110000+02	.48142127+01	DCR = 0.00189
4	.21480000+02	.67591441+01	
5	.26850000+02	.87366450+01	
6	.32220000+02	.10720590+02	
7	.37590000+02	.12719520+02	
8	.42960000+02	.14723384+02	
9	.48330000+02	.16741999+02	
10	.53700000+02	.18772583+02	
11	.59070000+02	.20808328+02	
12	.64440000+02	.22856739+02	
13	.69809999+02	.24911332+02	
14	.75179999+02	.26981318+02	
15	.80549999+02	.29052299+02	
16	.85920000+02	.31136083+02	
17	.91290000+02	.33222325+02	
18	.96660000+02	.35319666+02	
19	.10203000+03	.37417463+02	
20	.10740000+03	.39527701+02	
21	.11277000+03	.41641009+02	
22	.11814000+03	.43763936+02	
23	.12351000+03	.45891600+02	
24	.12888000+03	.48027113+02	
25	.13425000+03	.50167837+02	
26	.13962000+03	.52311296+02	
27	.14499000+03	.54464463+02	
28	.15036000+03	.56620806+02	
29	.15573000+03	.58781849+02	
30	.16110000+03	.60945718+02	
31	.16647000+03	.63112675+02	
32	.17184000+03	.65283478+02	
33	.17721000+03	.67454995+02	
34	.18258000+03	.69633939+02	
35	.18795000+03	.71811108+02	
36	.19332000+03	.73995925+02	
37	.19869000+03	.76184711+02	
38	.20406000+03	.78382043+02	
39	.20943000+03	.80589046+02	
40	.21480000+03	.82802028+02	
41	.22017000+03	.85026340+02	
42	.22554000+03	.87254579+02	
43	.23091000+03	.89493860+02	
44	.23628000+03	.91739651+02	
45	.24165000+03	.93990516+02	
46	.24702000+03	.96246378+02	
47	.25239000+03	.98508530+02	
48	.25776000+03	.10077140+03	
49	.26313000+03	.10304193+03	
50	.26850000+03	.10531535+03	
51	.27387000+03	.10759699+03	
52	.27924000+03	.109886948+03	
53	.28461000+03	.11218662+03	
54	.28998000+03	.11449390+03	

55	.29535000+03	.11680606+03
56	.30072000+03	.11911759+03
57	.30609000+03	.12143670+03
58	.31146000+03	.12375715+03
59	.31683000+03	.12607092+03
60	.32220000+03	.12838469+03
61	.32757000+03	.13070072+03
62	.33294000+03	.13301658+03
63	.33831000+03	.13533265+03
64	.34368000+03	.13764872+03
65	.34905000+03	.14016927+03
66	.35442000+03	.14253191+03
67	.35979000+03	.14490176+03
68	.36516000+03	.14726997+03

0 LINES NO SAMS AREA(KHXKMI) VARIANCE STD DEV

1	2	10.74	10.74	.72021	.84865
2	3	17.61	16.11	.26746	.51716
3	5	24.49	26.85	.11649	.34130
4	6	31.36	32.22	.07445	.27285
5	7	38.23	37.59	.05192	.22785
6	8	45.11	42.96	.03834	.19581
7	10	51.98	53.70	.02682	.16376
8	11	58.86	59.07	.02150	.14662
9	12	65.73	64.44	.01704	.13280
10	14	72.60	75.18	.01377	.11733
11	15	79.48	80.55	.01174	.10834
12	16	86.35	85.92	.01014	.10067
13	17	93.22	91.29	.00884	.09404
14	19	100.10	102.03	.00740	.08604
15	20	106.97	107.40	.00659	.08117
16	21	113.84	112.77	.00590	.07682
17	22	120.72	118.14	.00532	.07293
18	24	127.59	128.88	.00463	.06806
19	25	134.46	134.25	.00422	.06500
20	26	141.34	139.62	.00367	.06220
21	28	148.21	150.36	.00344	.05864
22	29	155.09	155.73	.00318	.05637
23	30	161.96	161.10	.00294	.05426
24	31	168.83	166.47	.00274	.05231
25	33	175.71	177.21	.00248	.04978
26	34	182.58	182.58	.00232	.04813
27	35	189.45	187.95	.00217	.04660
28	37	196.33	198.69	.00199	.04458
29	38	203.20	204.06	.00187	.04326
30	39	210.07	209.43	.00177	.04203
31	40	216.95	214.80	.00167	.04086
32	42	223.82	225.54	.00155	.03932
33	43	230.70	230.91	.00147	.03830
34	44	237.57	236.28	.00139	.03733
35	46	244.44	247.02	.00130	.03605
36	47	251.32	252.39	.00124	.03520
37	48	258.20	257.76	.00118	.03438
38	49	265.07	263.13	.00113	.03361
39	51	271.94	273.57	.00106	.03257



40	52278.81279.24	.00102	.03187
41	53285.68284.61	.00097	.03121
42	54292.56289.98	.00093	.03058
43	56299.43300.72	.00088	.02972
44	57306.30306.09	.00085	.02915
45	58313.18311.46	.00082	.02859
46	60320.05322.20	.00078	.02785
47	61326.93327.57	.00075	.02734
48	62333.80332.94	.00072	.02686
49	63340.67338.31	.00070	.02639
50	65347.55349.05	.00066	.02576

N



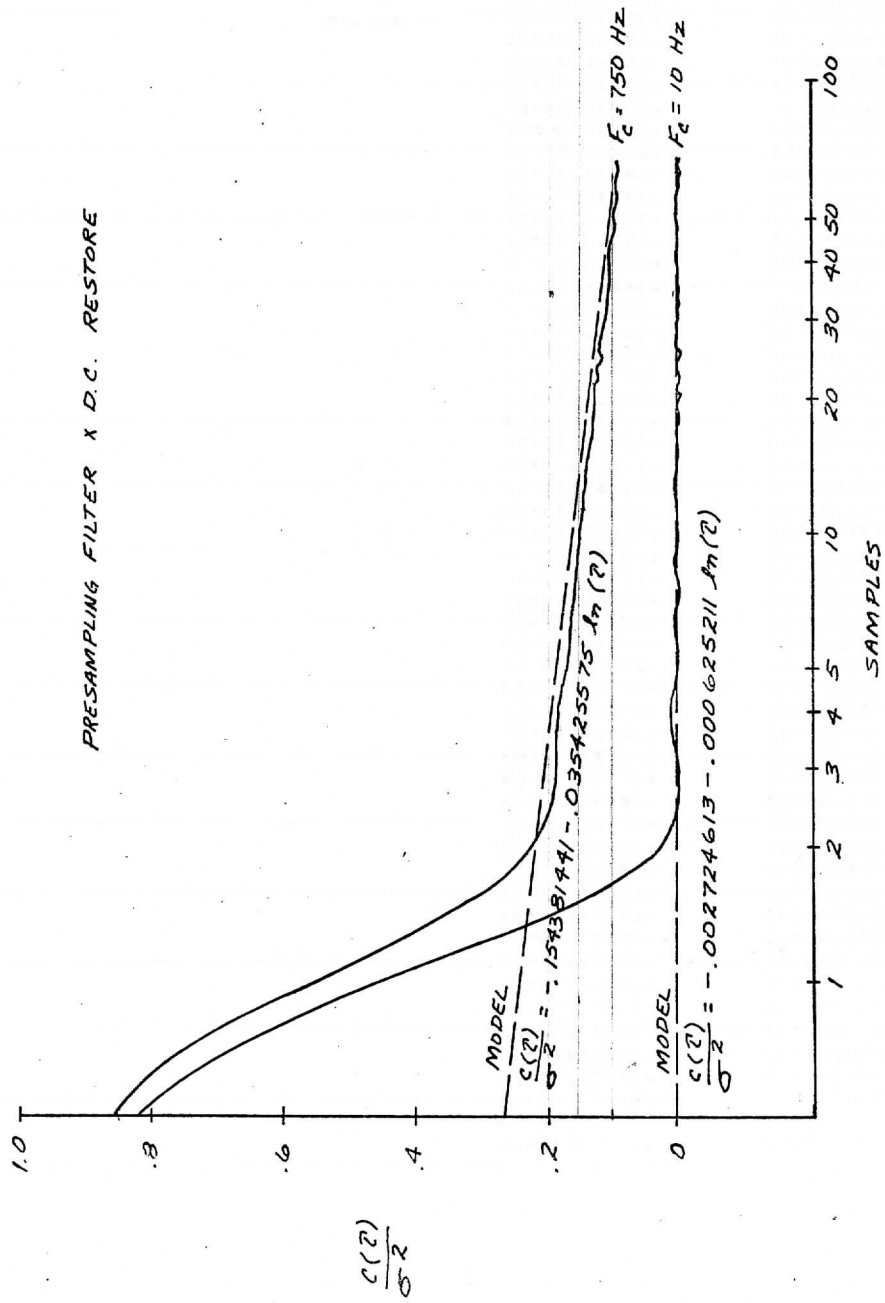


Figure 5

SAMPLE NO. KM VARIANCE OF MEAN

SAMPLE NO.	KM	VARIANCE OF MEAN	
1	.53700000+01	.10000000+01	WITH D. C. RESTORE F <sub>C</sub> = 750 Hz DCR = 0
2	.10740000+02	.31098800+01	
3	.16110000+02	.56543399+01	
4	.21480000+02	.85762398+01	
5	.26850000+02	.11868240+02	
6	.32220000+02	.15501799+02	
7	.37590000+02	.19467098+02	
8	.42960000+02	.23749717+02	
9	.48330000+02	.28347356+02	
10	.53700000+02	.33250575+02	
11	.59070000+02	.38147000+02	
12	.64440000+02	.43939045+02	
13	.69809999+02	.49715301+02	
14	.75179999+02	.55773634+02	
15	.80549999+02	.62104767+02	
16	.85920000+02	.68706697+02	
17	.91290000+02	.75573785+02	
18	.96660000+02	.82703712+02	
19	.10203000+03	.90087658+02	
20	.10740000+03	.97730503+02	
21	.11277000+03	.10562071+03	
22	.11814000+03	.11376293+03	
23	.12351000+03	.12214678+03	
24	.12888000+03	.13077605+03	
25	.13425000+03	.13964115+03	
26	.13962000+03	.14874273+03	
27	.14499000+03	.15807858+03	
28	.15036000+03	.16764573+03	
29	.15573000+03	.17743936+03	
30	.16110000+03	.18745831+03	
31	.16647000+03	.19769915+03	
32	.17184000+03	.20816083+03	
33	.17721000+03	.21883964+03	
34	.18258000+03	.22973735+03	
35	.18795000+03	.24084715+03	
36	.19332000+03	.25217247+03	
37	.19869000+03	.26371090+03	
38	.20406000+03	.27546268+03	
39	.20943000+03	.28742855+03	
40	.21480000+03	.29960339+03	
41	.22017000+03	.31198958+03	
42	.22554000+03	.32458105+03	
43	.23091000+03	.33737994+03	
44	.23628000+03	.35038329+03	
45	.24165000+03	.36358655+03	
46	.24702000+03	.37698979+03	
47	.25239000+03	.39059138+03	
48	.25776000+03	.40439341+03	
49	.26313000+03	.41837852+03	
50	.26850000+03	.43256321+03	
51	.27387000+03	.44694344+03	
52	.27924000+03	.46151924+03	
53	.28461000+03	.47528591+03	
54	.28998000+03	.48922813+03	

55	.29535000+03	.50639244+03
56	.30072000+03	.52172172+03
57	.30609000+03	.53723830+03
58	.31146000+03	.55293661+03
59	.31683000+03	.56861656+03
60	.32220000+03	.58433265+03
61	.32757000+03	.60113617+03
62	.33294000+03	.61756895+03
63	.33831000+03	.63417778+03
64	.34368000+03	.65095140+03
65	.34905000+03	.66791898+03
66	.35442000+03	.68505540+03
67	.35979000+03	.70235852+03
68	.36516000+03	.71985026+03

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 LINES NO SAMS AREA(KMXXKM) VARIANCE STD DEV  
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1	2	10.74	10.74	.77747	.88174
2	3	17.61	16.11	.31413	.56047
3	5	24.49	26.85	.15824	.39780
4	6	31.36	32.22	.10765	.32810
5	7	38.23	37.59	.07946	.28188
6	8	45.11	42.96	.06185	.24869
7	10	51.98	53.70	.04750	.21795
8	11	58.86	59.07	.03972	.19930
9	12	65.73	64.44	.03390	.18413
10	14	72.60	75.18	.02846	.16869
11	15	79.48	80.55	.02509	.15841
12	16	86.35	85.92	.02237	.14955
13	17	93.22	91.29	.02012	.14183
14	19	100.10	102.03	.01783	.13351
15	20	106.97	107.40	.01629	.12763
16	21	113.84	112.77	.01497	.12235
17	22	120.72	118.14	.01383	.11759
18	24	127.59	128.88	.01261	.11231
19	25	134.46	134.25	.01176	.10844
20	26	141.34	139.62	.01100	.10489
21	28	148.21	150.36	.01018	.10091
22	29	155.09	155.73	.00959	.09773
23	30	161.96	161.10	.00906	.09516
24	31	168.83	166.47	.00857	.09258
25	33	175.71	177.21	.00804	.08966
26	34	182.58	182.58	.00764	.08743
27	35	189.45	187.95	.00728	.08533
28	37	196.33	198.69	.00688	.08294
29	38	203.20	204.06	.00658	.08111
30	39	210.07	209.43	.00630	.07937
31	40	216.95	214.80	.00604	.07772
32	42	223.82	225.54	.00575	.07583
33	43	230.70	230.91	.00553	.07436
34	44	237.57	236.26	.00532	.07276
35	46	244.44	247.02	.00509	.07135
36	47	251.32	252.39	.00491	.07008
37	48	258.19	257.75	.00474	.06887
38	49	265.06	263.13	.00459	.06772
39	51	271.94	273.87	.00441	.06663

40	52278.81279.24	.00427	.56532
41	53285.68284.61	.00414	.56431
42	54292.56289.98	.00401	.56333
43	56299.43300.72	.00387	.56228
44	57306.30306.09	.00376	.56130
45	58313.18311.46	.00365	.56044
46	60320.05322.20	.00353	.55943
47	61326.93327.57	.00344	.55863
48	62333.80332.94	.00335	.55785
49	63340.67338.31	.00326	.55710
50	65347.55349.05	.00316	.55623
DCR	.00000		

FILE NO. KM VARIANCE OF MEAN

FILE NO.	KM	VARIANCE OF MEAN	
1	.53700000+01	.10000000+01	WITH D. C. RESTORE
2	.10740000+02	.28748200+01	F <sub>C</sub> = 10 Hz
3	.16110000+02	.47994840+01	DCR = 0
4	.21480000+02	.67356619+01	
5	.26850000+02	.86977676+01	
6	.32220000+02	.10667254+02	
7	.37590000+02	.12646484+02	
8	.42960000+02	.14629430+02	
9	.48330000+02	.16623838+02	
10	.53700000+02	.18627125+02	
11	.59070000+02	.20634511+02	
12	.64440000+02	.22648465+02	
13	.69809999+02	.24670719+02	
14	.75179999+02	.26700954+02	
15	.80549999+02	.28733751+02	
16	.85920000+02	.30772535+02	
17	.91290000+02	.32814815+02	
18	.96660000+02	.34862620+02	
19	1.02030000+03	.36909721+02	
20	1.07400000+03	.38966578+02	
21	1.12770000+03	.41022513+02	
22	1.18140000+03	.43087492+02	
23	1.23510000+03	.45151925+02	
24	1.28880000+03	.47224459+02	
25	1.34250000+03	.49296247+02	
26	1.39620000+03	.51371479+02	
27	1.44990000+03	.53450525+02	
28	1.50360000+03	.55532730+02	
29	1.55730000+03	.57614946+02	
30	1.61100000+03	.59698602+02	
31	1.66470000+03	.61782032+02	
32	1.71840000+03	.63866536+02	
33	1.77210000+03	.65949812+02	
34	1.82580000+03	.68036604+02	
35	1.87950000+03	.70120354+02	
36	1.93320000+03	.72207657+02	
37	1.98690000+03	.74297496+02	
38	2.04060000+03	.76392137+02	
39	2.09430000+03	.78494457+02	
40	2.14800000+03	.80599904+02	
41	2.20170000+03	.82713480+02	
42	2.25540000+03	.84829362+02	
43	2.30910000+03	.86952231+02	
44	2.36280000+03	.89080267+02	
45	2.41650000+03	.91209515+02	
46	2.47020000+03	.93341977+02	
47	2.52390000+03	.95477417+02	
48	2.57760000+03	.97611108+02	
49	2.63130000+03	.99749078+02	
50	2.68500000+03	1.0188858+03	
51	2.73870000+03	1.0403330+03	
52	2.79240000+03	1.0618551+03	
53	2.84610000+03	1.0834003+03	
54	2.89980000+03	1.1050200+03	

55	.29535000+03	.11266631+03
56	.30072000+03	.11482709+03
57	.30609000+03	.11699267+03
58	.31146000+03	.11915724+03
59	.31683000+03	.12132199+03
60	.32220000+03	.12348394+03
61	.32757000+03	.12567659+03
62	.33294000+03	.12785451+03
63	.33831000+03	.13004154+03
64	.34368000+03	.13222202+03
65	.34905000+03	.13440107+03
66	.35442000+03	.13658616+03
67	.35979000+03	.13877573+03
68	.36516000+03	.14096074+03

0 LINES NO SAMS AREA(KMXXN) VARIANCE STD DEV

1	2 10.74 10.74	.71870	.84776
2	3 17.61 16.11	.26664	.51637
3	5 24.49 26.85	.11597	.34054
4	6 31.36 32.22	.07408	.27217
5	7 38.23 37.59	.05162	.22720
6	8 45.11 42.96	.03810	.19519
7	10 51.98 53.70	.02661	.16313
8	11 58.86 59.07	.02132	.14600
9	12 65.73 64.44	.01746	.13220
10	14 72.60 75.18	.01362	.11672
11	15 79.48 80.55	.01161	.10775
12	16 86.35 85.97	.01002	.10009
13	17 93.22 91.29	.00873	.09346
14	19 100.10 102.03	.00730	.08546
15	20 106.97 107.40	.00649	.08059
16	21 113.84 112.77	.00581	.07625
17	22 120.72 118.14	.00524	.07236
18	24 127.59 128.88	.00455	.06749
19	25 134.46 134.25	.00415	.06443
20	26 141.34 139.62	.00380	.06164
21	28 148.21 150.36	.00337	.05808
22	29 155.09 155.73	.00311	.05580
23	30 161.96 161.10	.00288	.05370
24	31 168.83 166.47	.00268	.05176
25	33 175.71 177.21	.00242	.04922
26	34 182.58 182.58	.00226	.04758
27	35 189.45 187.95	.00212	.04604
28	37 196.33 198.69	.00194	.04403
29	38 203.20 204.06	.00182	.04271
30	39 210.07 209.43	.00172	.04148
31	40 216.95 214.80	.00162	.04031
32	42 223.82 226.59	.00151	.03977
33	43 230.70 230.91	.00140	.03975
34	44 237.57 238.28	.00135	.03879
35	46 244.44 247.02	.00126	.03550
36	47 251.32 252.39	.00120	.03465
37	48 258.19 257.78	.00115	.03374
38	49 265.06 263.13	.00109	.03326
39	51 271.93 273.97	.00103	.03202





[5] Calculations with DCR; without the bump

In this case we have only the first 29 bandpass filters from Table 3.

The procedure remains the same.  $\frac{C(\tau)}{\sigma^2}$  and the models are shown in Figure 6.

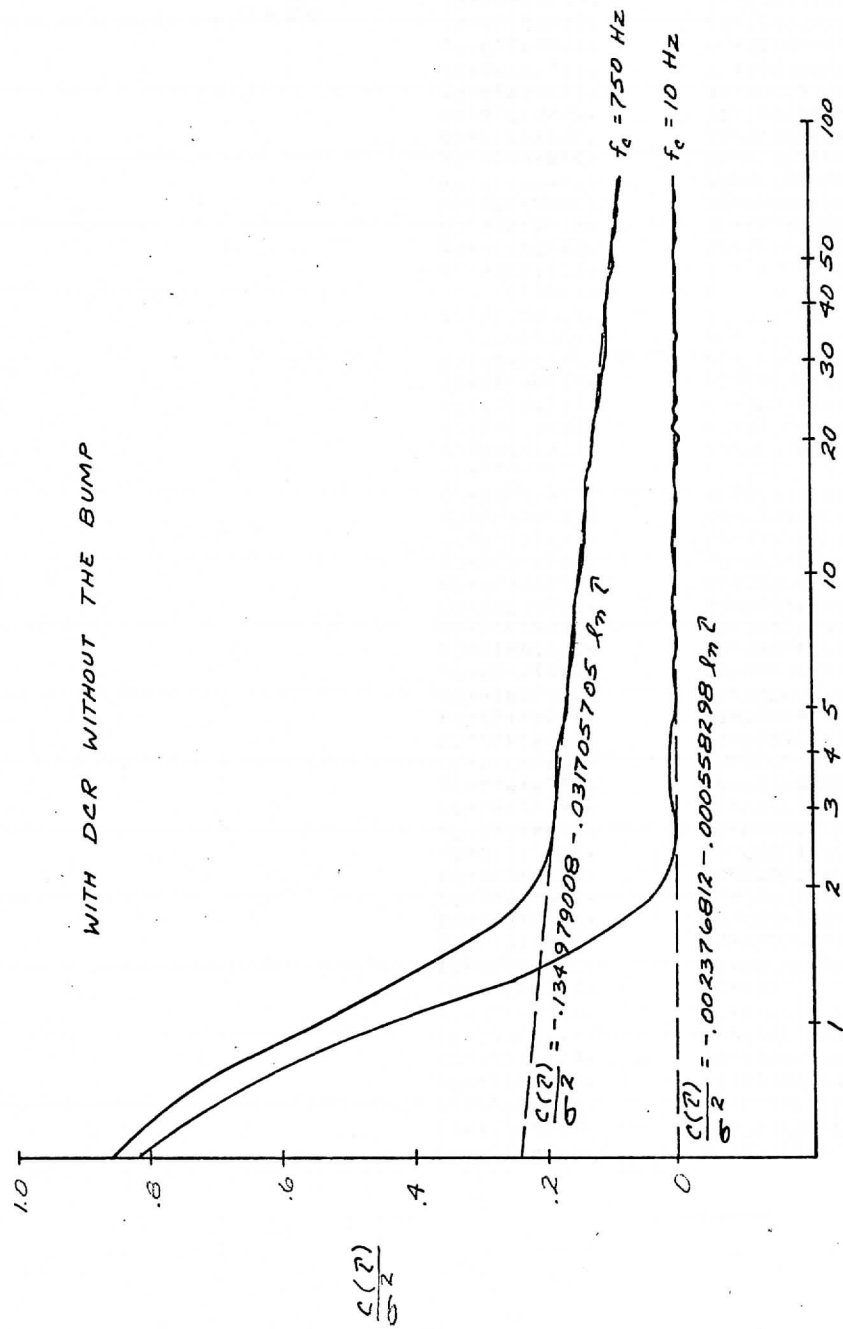
The models are:

$$\text{For } f_c = 750 \text{ Hz: } \frac{C(\tau)}{\sigma^2} = -.1350 - .03171n\tau$$

$$\text{For } f_c = 10 \text{ Hz: } \frac{C(\tau)}{\sigma^2} = -.0024 - .00061n\tau$$

A good agreement between the models and accurate calculations is seen in this case.

Table 5 shows the variance and standard deviation results for this case.



SAMPLES  
Figure 6

Table 5

AMPLE NO	KM	VARIANCE OF MEAN	
1	.53700000+01	.10009000+01	WITH D. C. RESTORE BUT WITHOUT THE BUMP
2	.10740000+02	.31077400+01	$F_c = 750 \text{ Hz}$
3	.16110000+02	.56462999+01	DCR = 0
4	.21480000+02	.65584197+01	
5	.26850000+02	.11836740+02	
6	.32220000+02	.15452659+02	
7	.37590000+02	.19396319+02	
8	.42960000+02	.23653278+02	
9	.48330000+02	.28221217+02	
10	.53700000+02	.33090676+02	
11	.59070000+02	.38451313+02	
12	.64440000+02	.43699368+02	
13	.69809999+02	.49429565+02	
14	.75179999+02	.55437779+02	
15	.80549999+02	.61714654+02	
16	.85920000+02	.68259388+02	
17	.91290000+02	.75062914+02	
18	.96660000+02	.82125920+02	
19	1.0203000+03	.89438986+02	
20	1.0740000+03	.97006771+02	
21	1.1277000+03	1.0481774+03	
22	1.1814000+03	1.1287654+03	
23	1.2351000+03	1.2117276+03	
24	1.2888000+03	1.2971023+03	
25	1.3425000+03	1.3847928+03	
26	1.3962000+03	1.4748060+03	
27	1.4499000+03	1.5671197+03	
28	1.5036000+03	1.6617040+03	
29	1.5573000+03	1.7585106+03	
30	1.6110000+03	1.8575281+03	
31	1.6647000+03	1.9587220+03	
32	1.7184000+03	2.0620818+03	
33	1.7721000+03	2.1675703+03	
34	1.8258000+03	2.2752054+03	
35	1.8795000+03	2.3849187+03	
36	1.9332000+03	2.4967447+03	
37	1.9869000+03	2.6106593+03	
38	2.0406000+03	2.7266649+03	
39	2.0943000+03	2.8447684+03	
40	2.1480000+03	2.9649193+03	
41	2.2017000+03	3.0871410+03	
42	2.2554000+03	3.2113727+03	
43	2.3091000+03	3.3376357+03	
44	2.3628000+03	3.4659008+03	
45	2.4165000+03	3.5961219+03	
46	2.4702000+03	3.7283002+03	
47	2.5239000+03	3.8624189+03	
48	2.5776000+03	3.9984820+03	
49	2.6313000+03	4.1363670+03	
50	2.6850000+03	4.2761874+03	
51	2.7387000+03	4.4179209+03	
52	2.7924000+03	4.5615678+03	
53	2.8461000+03	4.7070781+03	
54	2.8998000+03	4.8544724+03	

55	.24535000+03	.50037042+03
56	.30072000+03	.51547131+03
57	.30609000+03	.53075513+03
58	.31146000+03	.54621636+03
59	.31683000+03	.56185493+03
60	.32220000+03	.57767534+03
61	.32757000+03	.59367883+03
62	.33294000+03	.60985733+03
63	.33831000+03	.62620747+03
64	.34368000+03	.64272813+03
65	.34905000+03	.65941838+03
66	.35442000+03	.67628323+03
67	.35979000+03	.69332034+03
68	.36516000+03	.71052180+03

-----  
 LINES NO SAMS AREA(KMXXM) VARIANCE STD DEV

1	2	10.74	10.74	.77693	.88144
2	3	17.61	16.11	.31368	.56007
3	5	24.49	26.85	.15782	.39727
4	6	31.36	32.22	.10731	.32758
5	7	38.23	37.59	.07917	.28137
6	8	45.11	42.96	.06160	.24819
7	10	51.98	53.70	.04727	.21742
8	11	58.86	59.07	.03952	.19879
9	12	65.73	64.44	.03372	.18363
10	14	72.60	75.18	.02828	.16818
11	15	79.48	80.55	.02494	.15791
12	16	86.35	85.92	.02222	.14906
13	17	93.22	91.29	.01998	.14135
14	19	100.10	102.03	.01770	.13303
15	20	106.97	107.40	.01617	.12715
16	21	113.84	112.77	.01486	.12188
17	22	120.72	118.14	.01372	.11713
18	24	127.59	128.88	.01251	.11185
19	25	134.46	134.25	.01166	.10799
20	26	141.34	139.62	.01091	.10444
21	28	148.21	150.36	.01009	.10046
22	29	155.09	155.73	.00950	.09749
23	30	161.96	161.10	.00897	.09473
24	31	168.83	166.47	.00849	.09216
25	33	175.71	177.21	.00796	.08923
26	34	182.58	182.58	.00757	.08701
27	35	189.45	187.95	.00721	.08492
28	37	196.33	198.69	.00681	.08253
29	38	203.20	204.06	.00651	.08069
30	39	210.07	209.43	.00623	.07896
31	40	216.95	214.80	.00598	.07732
32	42	223.82	225.54	.00569	.07543
33	43	230.70	233.91	.00547	.07376
34	44	237.57	236.29	.00527	.07256
35	46	244.44	247.02	.00503	.07095
36	47	251.32	252.39	.00486	.06969
37	48	258.19	257.75	.00469	.06849
38	49	265.06	263.13	.00453	.06733
39	51	271.94	273.87	.00436	.06599

40	52278.81279.24	.00422	.06494
41	53285.68284.61	.00409	.06393
42	54292.56289.98	.00396	.06296
43	56299.43300.72	.00382	.06183
44	57306.30306.09	.00371	.06093
45	58313.18311.46	.00361	.06007
46	60320.05322.20	.00349	.05906
47	61326.93327.57	.00339	.05826
48	62333.80332.94	.00331	.05749
49	63340.67338.31	.00322	.05674
50	65347.55349.05	.00312	.05587

FILE NO. KM VARIANCE OF MEAN

1	.53700000+01	.10000000+01	
2	.10740000+02	.28746600+01	WITH D. C. RESTORE BUT WITHOUT
3	.16110000+02	.47787050+01	THE BUMP
4	.21430000+02	.67348028+01	$F_C = 10 \text{ Hz}$
5	.26850000+02	.86955688+01	DCR = 0
6	.32220000+02	.10663855+02	
7	.37590000+02	.12641606+02	
8	.42960000+02	.14622818+02	
9	.48330000+02	.16615229+02	
10	.53700000+02	.18616256+02	
11	.59070000+02	.20621118+02	
12	.64440000+02	.22632786+02	
13	.69809999+02	.24651489+02	
14	.75179999+02	.26678411+02	
15	.80549999+02	.28707630+02	
16	.85920000+02	.30742572+02	
17	.91290000+02	.32780748+02	
18	.96660000+02	.34824184+02	
19	.10203000+03	.36866651+02	
20	.10740000+03	.38918612+02	
21	.11277000+03	.40969385+02	
22	.11814000+03	.43028940+02	
23	.12351000+03	.45087685+02	
24	.12888000+03	.47154267+02	
25	.13425000+03	.49219838+02	
26	.13962000+03	.51288590+02	
27	.14499000+03	.53360893+02	
28	.15036000+03	.55436090+02	
29	.15573000+03	.57511036+02	
30	.16110000+03	.59587157+02	
31	.16647000+03	.61662788+02	
32	.17184000+03	.63739231+02	
33	.17721000+03	.65814179+02	
34	.18258000+03	.67892381+02	
35	.18795000+03	.69967277+02	
36	.19332000+03	.72045460+02	
37	.19869000+03	.74125916+02	
38	.20406000+03	.76210909+02	
39	.20943000+03	.78303321+02	
40	.21480000+03	.80398593+02	
41	.22017000+03	.82501733+02	
42	.22554000+03	.84606915+02	
43	.23091000+03	.86718822+02	
44	.23628000+03	.88835632+02	
45	.24165000+03	.90953391+02	
46	.24702000+03	.93074122+02	
47	.25239000+03	.95197529+02	
48	.25776000+03	.97319823+02	
49	.26313000+03	.99444973+02	
50	.26850000+03	.10157107+03	
51	.27387000+03	.10370273+03	
52	.27924000+03	.10584161+03	
53	.28461000+03	.10798284+03	
54	.28998000+03	.11013066+03	

55	.29535000+03	.11228085+03
56	.30072000+03	.11442726+03
57	.30609000+03	.11657820+03
58	.31146000+03	.11872786+03
59	.31683000+03	.12087744+03
60	.32220000+03	.12303397+03
61	.32757000+03	.12520043+03
62	.33294000+03	.12736790+03
63	.33831000+03	.12953537+03
64	.34368000+03	.13169773+03
65	.34905000+03	.13386005+03
66	.35442000+03	.13602815+03
67	.35979000+03	.13820045+03
68	.36516000+03	.14036794+03

0-LINE-S--NO--SABS--AREA(KPKRM)--VARIANCE--STD. DEV.

1	2 10.74	10.74	.71866	.84774
2	3 17.61	16.11	.26661	.51634
3	5 24.49	26.85	.11594	.34050
4	6 31.36	32.22	.07405	.27213
5	7 38.23	37.59	.05160	.22715
6	8 45.11	42.96	.03808	.19514
7	10 51.98	53.70	.02659	.16308
8	11 58.86	59.07	.02130	.14595
9	12 65.73	64.74	.01746	.13210
10	14 72.60	75.18	.01361	.11667
11	15 79.48	80.55	.01107	.10771
12	16 86.35	85.92	.01001	.10004
13	17 93.22	91.29	.00873	.09341
14	19 100.10	102.03	.00729	.08541
15	20 106.97	107.40	.00649	.08054
16	21 113.84	112.77	.00581	.07620
17	22 120.72	118.14	.00523	.07232
18	24 127.59	128.88	.00455	.06744
19	25 134.46	134.25	.00414	.06438
20	26 141.34	139.62	.00379	.06159
21	28 148.21	150.36	.00337	.05803
22	29 155.09	155.73	.00311	.05575
23	30 161.96	161.10	.00288	.05365
24	31 168.83	166.47	.00267	.05171
25	33 175.71	177.21	.00242	.04917
26	34 182.58	182.58	.00226	.04753
27	35 189.45	187.95	.00212	.04599
28	37 196.33	198.69	.00193	.04397
29	38 203.20	204.06	.00182	.04266
30	39 210.07	209.43	.00172	.04143
31	40 216.95	214.80	.00162	.04026
32	42 223.82	225.54	.00150	.03871
33	43 230.70	230.91	.00142	.03770
34	44 237.57	236.28	.00135	.03674
35	46 244.44	247.02	.00126	.03545
36	47 251.32	252.39	.00120	.03460
37	48 258.19	257.76	.00114	.03379
38	49 265.06	263.13	.00109	.03301
39	51 271.94	273.87	.00102	.03197



40	52278.81279.24	.00098	.03128
41	53285.68284.61	.00094	.03062
42	54292.56289.98	.00090	.02999
43	56299.43300.72	.00085	.02913
44	57306.30300.09	.00082	.02850
45	58313.18311.46	.00078	.02801
46	60320.05322.20	.00074	.02726
47	61326.93327.57	.00072	.02676
48	62333.80332.94	.00069	.02627
49	63340.67338.31	.00067	.02581
50	65347.55349.05	.00063	.02517

N



```

D=S1*S2*S3*S4*S5
T=1./D
A=RFAL(T)
B=AIMAG(T)
C=CARS(T)
FI=ATAN(B/A) * 180./3.14159265
2 FORMAT(7F10.5)
RETURN
END
FOR ISZ NOISTF
SUBROUTINE NOISTF(F,T1,T2,TNOISE)
TWOPI=2.*3.14159265
A=TWOPI*F*T1
B=TWOPI*F*T2
C=1.+B*B
D=1-2.*SIN(A)/(A*SQRT(C))+1./C
TNOISE=SQRT(D)
RETURN
END
IXQT
0.70557
0.67745 -0.94007
0.67745 0.94007
0.54117 -1.82565
0.54117 1.82565
0.026 17450.
IFIN

```

## APPENDIX B

I. Calculation of the Autocovariance Function  $C(\tau)$ 

$$C(\tau) = \int_0^{\infty} T(f)P(f)\cos 2\pi f\tau df$$

$P(f)$  = noise power spectral density in the detector

$T(f)$  = power transfer function of the electronics following the detector

Special case

$$P(f) = K\left(1 + \frac{c}{f}\right)$$

$$T(f) = \begin{cases} 1 & \text{for } f_1 \leq f \leq f_2 \\ 0 & \text{for } f < f_1 \text{ or } f > f_2 \end{cases}$$

$$C(\tau) = \int_{f_1}^{f_2} P(f)\cos 2\pi f\tau df$$

$$C(0) = \sigma^2 = \int_{f_1}^{f_2} P(f)df = K(f_2 - f_1 + f_c \ln \frac{f_2}{f_1}) = K\Delta f_N$$

$$\Rightarrow K = \frac{\sigma^2}{\Delta f_N}$$

$$C(\tau) = \frac{\sigma^2}{\Delta f_N} \int_{f_1}^{f_2} \left(1 + \frac{f_c}{f}\right) \cos 2\pi f\tau df$$

Model for special case

$$C(\tau) = \sigma^2 \frac{f_c}{\Delta f_N} [-\gamma - \ln 2\pi f_{\min} \tau] \quad \text{for } 2\pi f_{\max} \tau > 1$$

where  $\gamma = .577216$

For  $\tau$  in seconds,  $f$ , in Hz we have

$$\frac{C(\tau)}{\sigma^2} = a - b \ln \tau,$$

where  $a = b(-\ln 2\pi f_{\min} - \gamma)$ ,  $b = \frac{f_c}{\Delta f_N}$ ,  $\Delta f_N = f_{\max} - f_{\min} + f_c \ln \frac{f_{\max}}{f_{\min}}$

General Case solved by series

$$T(f) = \sum_{k=1}^N W_k T_k(f)$$

$$\text{where } T_k(f) = \begin{cases} 1 & f_{\min_k} \leq f \leq f_{\max_k} \\ 0 & f \leq f_{\min_k} \text{ or } f > f_{\max_k} \end{cases}$$

can be used to approximate a bandpass filter with realistic rolloff or the combined effects of bandpass and DC restore filters.

For the series approximation we have

$$C(\tau) = \sum_{k=1}^N W_k \int_{f_{1,k}}^{f_{2,k}} P(f) \cos 2\pi f \tau df, \text{ and}$$

$$\sigma^2 = \sum_{k=1}^N W_k \int_{f_{1,k}}^{f_{2,k}} P(f) df = K \sum_{k=1}^N W_k [f_{2,k} - f_{1,k} + f_c \ln \frac{f_{2,k}}{f_{1,k}}]$$

Thus

$$\frac{C(\tau)}{\sigma^2} = \frac{\sum_{k=1}^N W_k \int_{f_{1,k}}^{f_{2,k}} (1 + \frac{f_c}{f}) \cos 2\pi f \tau df}{\sum_{k=1}^N W_k [f_{2,k} - f_{1,k} + f_c \ln \frac{f_{2,k}}{f_{1,k}}]}$$

model

$$\int_{f_{1,k}}^{f_{2,k}} (1 + \frac{f_c}{f}) \cos 2\pi f \tau df \approx f_c [-\gamma - \ln 2\pi f_{1,k}] - \ln \tau$$

$$\text{where } \Delta f_{N,k} = f_{2,k} - f_{1,k} + f_c \ln \frac{f_{2,k}}{f_{1,k}}$$

$$\frac{C(\tau)}{\sigma^2} \approx \frac{\sum_{k=1}^N W_k f_c [-\gamma - \ln 2\pi f_{1,k}] - \ln \tau}{\sum_{k=1}^N W_k \Delta f_{N,k}}$$

$$\frac{C(\tau)}{\sigma^2} = \left[ \frac{\sum_{k=1}^N W_k f_c [-\gamma - \ln 2\pi f_{1,k}]}{\sum_{k=1}^N W_k \Delta f_{N,k}} \right] - \left[ \frac{\sum_{k=1}^N W_k f_c}{\sum_{k=1}^N W_k \Delta f_{N,k}} \right] \ln \tau$$

$\approx \bar{a} - \bar{b} \ln \tau$ , where

$$a = \frac{\sum_{k=1}^N W_k f_c [-\gamma - \ln 2\pi f_{1,k}]}{\sum_{k=1}^N W_k \Delta f_{N,k}} \quad \text{and} \quad b = \frac{\sum_{k=1}^N W_k f_c}{\sum_{k=1}^N W_k \Delta f_{N,k}}$$

Calculation of autocovariance function:

$$\frac{C(\tau)}{\sigma^2} = \bar{a} - \bar{b} \ln \tau, \text{ where}$$

$$\bar{a} = \frac{\sum_{k=1}^N W_k f_c [-\gamma - \ln 2\pi f_{1,k}]}{\sum_{k=1}^N W_k \Delta f_{N,k}} \quad \text{SUM3}$$

$$\bar{b} = \frac{\sum_{k=1}^N W_k f_c}{\sum_{k=1}^N W_k \Delta f_{N,k}} \quad \begin{array}{l} \text{SUM1} \\ \text{SUM2} \end{array}$$

GAMMA = .577216

```

READ FC
DO 10 I=1, N
READ101, FMIN, FMAX, W
101 FORMAT
DELF = FMAX - FMIN + FC*ALOG(FMAX/FMIN)
TOPIFM=2.*3.14159265*FMIN
SUM1=SUM1+W*FC
SUM2=SUM2+W*DELF
SUM3=SUM3+W*FC*[-GAMA-ALOG(TOPIFM/
10 CONTINUE
A=SUM3/SUM2
B=SUM1/SUM2
PRINT102A, B
102 FORMAT('C(TAU)=,F10.5, '+' ,F10.5, '*ALOG(TAU) ')

```



```

COV(I)=0.
DO 100 L=1,K
100 COV(I)=COV(I)+A(L,I)*W(L)
110 COV(I)=COV(I)/SUMSI2
PRINT 2,DELTA
2 FORMAT('DELTA=',F20.9//WEIGHTS ',(10F10.6))
PRINT 4
4 FORMAT('POSITION TAU COV')
DO 102 I=1,N
102 PRINT 3,I,TAU(I),COV(I)
3 FORMAT('I ',I,' TAU ',I2.7)
PUNCH,(TAU(I),COV(I)),I=1,N)
GO TO 1
33 STOP
END
call graph(100,TAU,COV,4, LABEL)
call graph
*FOR,ISZ COVAR,COVAR
SUBROUTINE COVAR(N,K,I,COVAR,TAU)
DIMENSION TAU(N),COVAR(K,N)
COMMON/FF/FC,FMIN,FMAX,DELT
TWOPI=2.*3.14159265
DO 10 KK=1,N
TAU(KK)=FLOAT(KK)*DELT
TWOPII= TWOPI*TAU(KK)
X1=TWOPII*FMAX
X2=TWOPII*FMIN
CIX1=FCI(X1)
CIX2=FCI(X2)
ROMB=CIX1-CIX2
ROMB=(SIN(TWOPII*FMAX)-SIN(TWOPII*FMIN))/TWOPII+FC*ROMB
C COVAR IS COVARIANCE FUNCTION C(TAU)/SIGMA**2
COVAR(I,KK)=ROMB
10 CONTINUE
PRINT 72,(KK,TAU(KK),COVAR(I,KK),KK=1,N)
72 FORMAT('I H ',2(' NO K TAU C(TAU)/SIGMA**2'))//
1 (2( I9,F10.6,5X,E12.6)))
RETURN
END
*FOR,ISZ FCI,FCI
FUNCTION FCI(X)
DIMENSION CIC1(21),CIC2(21)
DIMENSION CIA(51),CI(151),CIB(81)
COMMON/CC/CIA,CI,CIB,CIC1,CIC2
GAMA=0.57722
IF(X.GT.10.)GO TO 4
IF(X.GE..50) GO TO 2
IX=INT(X*100.)+1
D=X-(IX-1)*0.01
Y=CIA(IX)+D*(CIA(IX+1)-CIA(IX))
FCI =Y*X*X+GAMA+ALOG(X)
RETURN
2 IF(X.GT.2.00) GO TO 3
IX=INT((X-.5)*100.)+1
D=X-0.5-(IX-1)*0.01
FCI=CIB(IX)+D*(CIB(IX+1)-CIB(IX))
RETURN
3 IX=INT((X-2.0)*10.)+1

```



D=X-2.-(IX-1)*0.1							
FCI=CIB(IX)+D*(CIB(IX+1)-CIB(IX))							
RETURN							
4 X1=1./X							
IX=INT(X1*200.)+1							
D=X1-(IX-1)/200.							
XFX=CIC1(IX)+D*(CIC1(IX+1)-CIC1(IX))							
X2G=CIC2(IX)+D*(CIC2(IX+1)-CIC2(IX))							
FCI=XFX*SIN(X)/X-X2G*COS(X)/(X*X)							
RETURN							
END							
MAP,I							
XQT							
-.25000	-.25000	-.25000	-.24999	-.24998	-.24997	-.24996	-.24995
-.24993	-.24992	-.24990	-.24987	-.24985	-.24982	-.24980	-.24977
-.24973	-.24970	-.24966	-.24962	-.24958	-.24954	-.24950	-.24945
-.24940	-.24935	-.24930	-.24924	-.24918	-.24913	-.24906	-.24900
-.24893	-.24887	-.24880	-.24873	-.24865	-.24858	-.24850	-.24842
-.24834	-.24826	-.24817	-.24808	-.24799	-.24790	-.24781	-.24771
-.24761	-.24751	-.24741					
-.17778	-.16045	-.14355	-.12707	-.11099	-.09530	-.07999	-0.06504
-.05044	-.03619	-.02227	-.00867	.00460	.1758	.03026	.04265
.05476	.06659	.07816	.08946	.10051	.11131	.12187	.13220
.14230	.15216	.16181	.17124	.18046	.18947	.19828	.20680
.21530	.22353	.23157	.23942	.24710	.25460	.26192	.26908
.27607	.28289	.28956	.29606	.30242	.30861	.31466	.32056
.32632	.33193	.33740	.34274	.34794	.35300	.35793	.36274
.36741	.37196	.37639	.38069	.38487	.38894	.39289	.39672
.40044	.40405	.40754	.41093	.41421	.41739	.42046	.42343
.42629	.42906	.43173	.43430	.43678	.43916	.44144	.44364
.44574	.44775	.44968	.45151	.45326	.45493	.45651	.45800
.45942	.46075	.46201	.46318	.46428	.46530	.46624	.46711
.46790	.46862	.46927	.46985	.47035	.47079	.47116	.47146
.47169	.47186	.47196	.47200	.47197	.47188	.47173	.47152
.47125	.47091	.47052	.47007	.46956	.46900	.46838	.46770
.46697	.46618	.46535	.46446	.46351	.46252	.46148	.46038
.45924	.45805	.45681	.45553	.45419	.45282	.45139	.44992
.44841	.44685	.44526	.44362	.44194	.44022	.43846	.43665
.43481	.43293	.43102	.42906	.42707	.42504	.42298	
.42298	.40051	.37507	.34718	.31729	.28587	.25334	.22008
.18649	.15290	.11963	.08699	.05526	.02468	-.00452	-.03213
-.05797	-.08190	-.10378	-.12350	-.14098	-.15617	-.16901	-.17951
-.18766	-.19349	-.19705	-.19839	-.19760	-.19478	-.19003	-.18348
-.17525	-.16551	-.15439	-.14205	-.12867	-.11441	-.09944	-.08393
-.06806	-.05198	-.03587	-.01989	-.00418	.01110	.02582	.03985
.05308	.06539	.07670	.08691	.09596	.10379	.11036	.11563
.11960	.12225	.12359	.12364	.12243	.12002	.11644	.11177
.10607	.09943	.09193	.08368	.07476	.06528	.05535	.04507
.03455	.02391	.01325	.00268	-.00770	-.01780	-.02752	-.03676
-.04546							
1.00000	1.00000	.99995	.99985	.99980	.99940	.99955	.99866
.99920	.99762	.99876	.99630	.99822	.99469	.99758	.99282
.99686	.99069	.99604	.98831	.99514	.98568	.99415	.98382
.99308	.97976	.99153	.97649	.99071	.97303	.98941	.96030
.98804	.96557	.98660	.96160	.98510	.95748	.98353	.95320
.98191	.94885						

C3

750. 0.00000716197

10.	0.026	17450.	
0.2	48000.	0.0025	C4
0.54	44500.	0.0025	
0.77	41000.	0.005	
1.1	39500.	0.01	
1.52	37000.	0.02	
2.2	34500.	0.04	
2.82	30300.	0.04	
3.47	28000.	0.04	
3.88	26500.	0.04	
4.3	25000.	0.04	
4.7	23750.	0.04	
5.03	22500.	0.04	
5.55	21500.	0.04	
5.9	20200.	0.04	
6.3	19300.	0.04	
5.0	18000.	0.04	
7.2	17500.	0.04	
7.6	16500.	0.04	
7.9	15400.	0.04	
8.3	14200.	0.04	
8.8	13200.	0.04	
9.3	12000.	0.04	
9.8	11000.	0.04	
10.2	9500.	0.04	
11.0	8500.	0.04	
11.5	7300.	0.04	
12.0	5900.	0.04	
12.7	4000.	0.04	
13.5	1900.	0.04	

FIN



750.	0.026	17450.	
0.3	48000.	0.0025	D2
0.54	44500.	0.0025	
0.77	41000.	0.005	
1.1	39500.	0.01	
1.52	37000.	0.02	
2.2	34000.	0.04	
2.82	30300.	0.04	
3.47	28000.	0.04	
3.88	26500.	0.04	
4.3	25000.	0.04	
4.7	23750.	0.04	
5.03	22500.	0.04	
5.55	21500.	0.04	
5.9	20200.	0.04	
6.3	19300.	0.04	
6.8	18500.	0.04	
7.2	17500.	0.04	
7.6	16500.	0.04	
7.9	15400.	0.04	
8.3	14200.	0.04	
8.8	13200.	0.04	
9.3	12000.	0.04	
9.8	11000.	0.04	
10.3	9800.	0.04	
11.0	8500.	0.04	
11.5	7300.	0.04	
12.0	5900.	0.04	
12.7	4000.	0.04	
13.5	1900.	0.04	
29			
10.	0.026	17450.	
0.3	48000.	0.0025	
0.54	44500.	0.0025	
0.77	41000.	0.005	
1.1	39500.	0.01	
1.52	37000.	0.02	
2.2	34000.	0.04	
2.82	30300.	0.04	
3.47	28000.	0.04	
3.88	26500.	0.04	
4.3	25000.	0.04	
4.7	23750.	0.04	
5.03	22500.	0.04	
5.55	21500.	0.04	
5.9	20200.	0.04	
6.3	19300.	0.04	
6.8	18500.	0.04	
7.2	17500.	0.04	
7.6	16500.	0.04	
7.9	15400.	0.04	
8.3	14200.	0.04	
8.8	13200.	0.04	
9.3	12000.	0.04	
9.8	11000.	0.04	
10.3	9800.	0.04	

11.0	8500.	0.04	
11.5	7300.	0.04	D3
12.0	5900.	0.04	
12.7	4000.	0.04	
13.5	1900.	0.04	
FIN			

APPENDIX E

\*RUN SUBHASHSSEC,4623,9000118597,2M,100/1000  
 \*FOR,ZSI RTSPAC

EI

```

DIMENSION COV(136),COVS(68),V(68),D(68)
DIMENSION TAU(136)
100 READ(-,-,END=33)DCR
PRINT 900,DCR
900 FORMAT(1H1,' DCR=',F10.5)
READ(-,-) ((TAU(I),COV(I)),I=1,136)
COVS(1)=1
DO 300 I=2,68
IJ=2*(I-1)
300 COVS(I)=(COV(IJ )-DCR)/(I.-DCR)
DO 400 I=1,68
V(I)=0.
DO 400 J=1,I
DO 400 K=1,I
IJ=1+IABS(J-K)
400 V(I)=V(I)+COV(IJ)
PRINT 2
2 FORMAT('1SAMPLE NO KM VARIANCE OF MEAN'//)
DO 5 I=1,68
D(I)=FLOAT(I)*2.9992
5 PRINT 1,I,D(I),V(I)
1 FORMAT(I6,=15.8,E20.8)
PRINT 3
3 FORMAT('10NO LINES NO SAMS AREA(KMxKM) VARIANCE STD DEV'//)
DO 6 NL=1,50
SEE NOTE BELOW.
NTOT=NL*NS
VA=FLOAT(NL)*V(NS)/FLOAT(NTOT**2)
DV=35.8*(.2+FLOAT(NL-1)*.192)
STDV=SQRT(VA)
6 PRINT 4,NL,NS,DV,D(NS),VA,STDV
4 FORMAT(I6,=18.2F6.2,F9.5,F9.5)
GO TO 100
33 STOP
END

```

\*MAP,IX  
 \*XOT  
 0.0

7.1620E-06	8.5572E-01	1.4324E-05	5.5387E-01	2.1486E-05
3.1622E-01	2.8648E-03	2.1541E-01	3.5810E-05	1.9136E-01
4.2972E-05	1.8678E-01	5.0134E-05	1.8541E-01	5.7296E-05
1.8310E-01	6.4458E-05	1.7663E-01	7.1620E-05	1.6880E-01
7.8782E-05	1.6504E-01	8.5944E-05	1.6387E-01	9.3106E-05
1.6048E-01	1.0027E-04	1.5665E-01	1.0743E-04	1.5568E-01
1.1459E-04	1.5549E-01	1.2175E-04	1.5391E-01	1.2892E-04
1.5076E-01	1.3608E-04	1.4727E-01	1.4324E-04	1.4559E-01
1.5040E-04	1.4521E-01	1.5756E-04	1.4371E-01	1.6473E-04
1.4166E-01	1.7189E-04	1.4108E-01	1.7905E-04	1.4095E-01
1.8621E-04	1.3899E-01	1.9337E-04	1.3584E-01	2.0054E-04
1.3434E-01	2.0770E-04	1.3441E-01	2.1486E-04	1.3343E-01
2.2202E-04	1.3136E-01	2.2918E-04	1.3040E-01	2.3635E-04
1.3031E-01	2.4351E-04	1.2924E-01	2.5067E-04	1.2685E-01
2.5783E-04	1.2503E-01	2.6499E-04	1.2597E-01	2.7215E-04
1.2736E-01	2.7932E-04	1.2503E-01	2.8648E-04	1.2159E-01

NOTE: Use this with 0.192mr FOV - NS=INT((.192+FLOAT(NL-1)\*.192)/.0838+.5)

Use this with 0.384mr FOV - NS=INT((.384+FLOAT(NL-1)\*.384)/.0838+.5)

2.9364E-04	1.2210E-01	3.0080E-04	1.2392E-01	3.0796E-04
1.2201E-01	3.1513E-04	1.1871E-01	3.2229E-04	1.1871E-01
3.2945E-04	1.2062E-01	3.3661E-04	1.1955E-01	3.4377E-04
1.1581E-01	3.5094E-04	1.1445E-01	3.5810E-04	1.1613E-01
3.6526E-04	1.1659E-01	3.7242E-04	1.1503E-01	3.7958E-04
1.1402E-01	3.8675E-04	1.1353E-01	3.9391E-04	1.1226E-01
4.0107E-04	1.1112E-01	4.0823E-04	1.1097E-01	4.1539E-04
1.1054E-01	4.2256E-04	1.0933E-01	4.2972E-04	1.0882E-01
4.3688E-04	1.0914E-01	4.4404E-04	1.0830E-01	4.5120E-04
1.0657E-01	4.5837E-04	1.0643E-01	4.6553E-04	1.0755E-01
4.7269E-04	1.0734E-01	4.7985E-04	1.0543E-01	4.8701E-04
1.0391E-01	4.9418E-04	1.0427E-01	5.0134E-04	1.0563E-01
5.0850E-04	1.0581E-01	5.1566E-04	1.0445E-01	5.2282E-04
1.0356E-01	5.2999E-04	1.0457E-01	5.3715E-04	1.0589E-01
5.4431E-04	1.0490E-01	5.5147E-04	1.0259E-01	5.5863E-04
1.0237E-01	5.6580E-04	1.0384E-01	5.7296E-04	1.0354E-01
5.8012E-04	1.0150E-01	5.8728E-04	1.0051E-01	5.9444E-04
1.0102E-01	6.0161E-04	1.0157E-01	6.0877E-04	1.0124E-01
6.1593E-04	1.0010E-01	6.2309E-04	9.8789E-02	6.3025E-04
9.7802E-02	6.3742E-04	9.7419E-02	6.4458E-04	9.7858E-02
6.5174E-04	9.8220E-02	6.5890E-04	9.7030E-02	6.6606E-04
9.5019E-02	6.7323E-04	9.4570E-02	6.8039E-04	9.5861E-02
6.8755E-04	9.6401E-02	6.9471E-04	9.5140E-02	7.0187E-04
9.4127E-02	7.0904E-04	9.4740E-02	7.1620E-04	9.5642E-02
7.2336E-04	9.5850E-02	7.3052E-04	9.5689E-02	7.3768E-04
9.4765E-02	7.4484E-04	9.3172E-02	7.5201E-04	9.2819E-02
7.5917E-04	9.4188E-02	7.6633E-04	9.4475E-02	7.7349E-04
9.1890E-02	7.8065E-04	8.8975E-02	7.8782E-04	8.8865E-02
7.9498E-04	9.0700E-02	8.0214E-04	9.1530E-02	8.0930E-04
9.0473E-02	8.1646E-04	8.8715E-02	8.2363E-04	8.7866E-02
8.3079E-04	8.8666E-02	8.3795E-04	9.0076E-02	8.4511E-04
9.0921E-02	8.5227E-04	9.1470E-02	8.5944E-04	9.1559E-02
8.6660E-04	9.0061E-02	8.7376E-04	8.7481E-02	8.8092E-04
8.5862E-02	8.8808E-04	8.5854E-02	8.9525E-04	8.6035E-02
9.0241E-04	8.5237E-02	9.0957E-04	8.4353E-02	9.1673E-04
8.4816E-02	9.2389E-04	8.6239E-02	9.3106E-04	8.7278E-02
9.3822E-04	8.7334E-02	9.4538E-04	8.6176E-02	9.5254E-04
8.3992E-02	9.5970E-04	8.2156E-02	9.6687E-04	8.2029E-02
9.7403E-04	8.3238E-02			
).0				
7.1620E-06	8.1727E-01	1.4324E-05	4.3733E-01	2.1486E-05
1.4294E-01	2.8648E-05	2.4793E-02	3.5810E-05	3.7037E-03
4.2972E-05	5.6255E-03	5.0134E-05	1.0265E-02	5.7296E-05
1.2833E-02	6.4458E-05	9.3368E-03	7.1620E-05	3.5635E-03
7.8782E-05	2.6808E-03	8.5944E-05	4.7306E-03	9.3106E-05
3.6297E-03	1.0027E-04	1.7307E-03	1.0743E-04	3.1885E-03
1.1459E-04	5.5995E-03	1.2175E-04	6.0208E-03	1.2892E-04
4.3086E-03	1.3608E-04	2.0232E-03	1.4324E-04	1.9172E-03
1.5040E-04	3.4507E-03	1.5756E-04	3.4030E-03	1.6473E-04
2.5326E-03	1.7189E-04	3.5182E-03	1.7905E-04	5.0217E-03
1.8621E-04	4.1090E-03	1.9337E-04	1.6336E-03	2.0054E-04
1.1486E-03	2.0770E-04	2.7105E-03	2.1486E-04	2.8621E-03
2.2202E-04	1.5472E-03	2.2918E-04	1.6165E-03	2.3635E-04
2.7742E-03	2.4351E-04	2.6316E-03	2.5067E-04	7.1105E-04
2.5783E-04	-4.8425E-04	2.6499E-04	1.9583E-03	2.7215E-04
4.7467E-03	2.7932E-04	2.8049E-03	2.8648E-04	-5.9341E-04

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2.9364E-04	1.0900E-03	3.0080E-04	4.3910E-03	3.0796E-04
2.9075E-03	3.1513E-04	-4.0556E-04	3.2229E-04	5.3694E-04
3.2946E-04	3.9196E-03	3.3661E-04	3.4443E-03	3.4377E-04
-5.0551E-04	3.5094E-04	-1.4073E-03	3.5510E-04	1.5907E-03
3.6526E-04	2.9922E-03	3.7242E-04	1.7753E-03	3.7958E-04
1.2851E-03	3.8675E-04	1.4478E-03	3.9391E-04	5.7595E-04
4.0107E-04	-1.2618E-04	4.0823E-04	4.1404E-04	4.1539E-04
5.8024E-04	4.2256E-04	-2.8010E-03	4.2972E-04	-2.4527E-04
4.3688E-04	8.3810E-04	4.4404E-04	4.0549E-04	4.5120E-04
-1.1773E-03	4.5837E-04	-7.4392E-04	4.6553E-04	1.3022E-03
4.7269E-04	1.6269E-03	4.7985E-04	-2.5333E-04	4.3701E-04
-1.6530E-03	4.9018E-04	-6.2374E-04	5.0134E-04	1.6450E-03
5.0650E-04	2.3850E-03	5.1566E-04	1.1367E-03	5.2232E-04
4.8470E-04	5.2999E-04	2.2691E-03	5.3715E-04	4.4615E-03
5.4431E-04	3.7095E-03	5.5147E-04	1.2223E-03	5.5863E-04
1.4312E-03	5.6580E-04	3.8327E-03	5.7296E-04	3.9340E-03
5.8017E-04	1.7948E-03	5.8728E-04	1.0214E-03	5.9444E-04
2.1833E-03	6.0161E-04	3.3632E-03	6.0877E-04	3.4231E-03
6.1593E-04	2.4512E-03	6.2309E-04	1.2574E-03	6.3025E-04
4.7511E-04	6.3742E-04	4.5264E-04	6.4458E-04	1.4861E-03
6.5174E-04	2.4137E-03	6.5890E-04	1.3374E-03	6.6606E-04
-8.0899E-04	6.7323E-04	-9.5612E-04	6.8039E-04	1.1275E-03
6.8755E-04	2.2285E-03	6.9471E-04	9.8995E-04	7.0187E-04
7.2760E-05	7.0904E-04	1.2455E-03	7.1620E-04	2.7822E-03
7.2336E-04	3.4266E-03	7.3052E-04	3.6132E-03	7.3768E-04
2.8307E-03	7.4484E-04	1.1708E-03	7.5201E-04	1.1187E-03
7.5917E-04	3.2969E-03	7.6633E-04	4.0915E-03	7.7349E-04
1.1860E-03	7.8065E-04	-2.1523E-03	7.8782E-04	-1.8920E-03
7.9498E-04	8.3565E-04	8.0214E-04	2.2671E-03	8.0930E-04
1.2662E-03	8.1646E-04	-6.3993E-04	8.2363E-04	-1.3900E-03
8.5079E-04	-3.7847E-05	8.3795E-04	2.0827E-03	8.4511E-04
3.4696E-03	8.5227E-04	4.4976E-03	8.5944E-04	4.9711E-03
8.6660E-04	3.4284E-03	8.7376E-04	5.0279E-04	8.8092E-04
-1.1856E-03	8.8808E-04	-3.1250E-04	8.9525E-04	-2.1847E-04
9.0241E-04	-9.1144E-04	9.0957E-04	-1.7374E-03	9.1673E-04
-8.4661E-04	9.2389E-04	1.2692E-03	9.3106E-04	2.8919E-03
9.3822E-04	3.2695E-03	9.4538E-04	2.1072E-03	9.5254E-04
-3.6835E-04	9.5970E-04	-2.4096E-03	9.6687E-04	-2.2737E-03
9.7403E-04	-4.4603E-04			

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